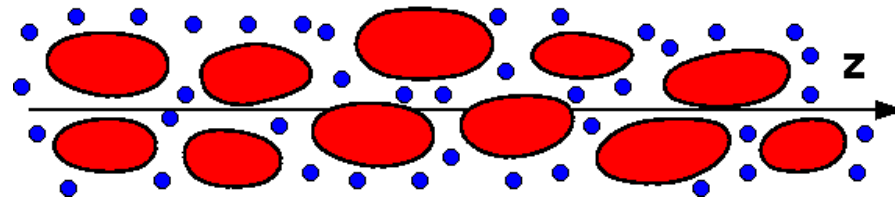


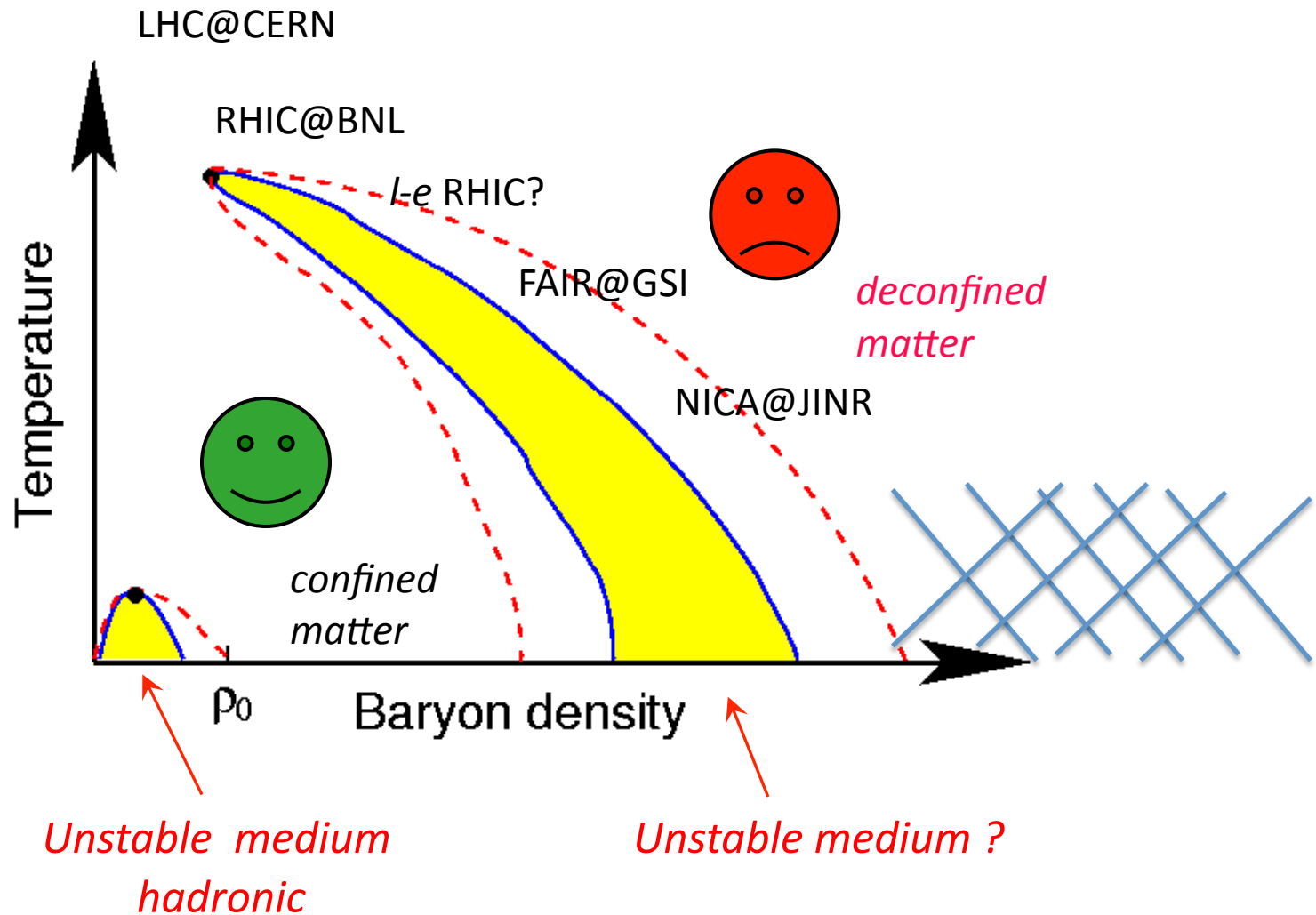
International Workshop on  
*In-Medium Effects in Hadronic and Partonic Systems*  
Oberurgl, Austria, 21 – 25 February 2011

*Spinodal phase decomposition  
with dissipative fluid dynamics*

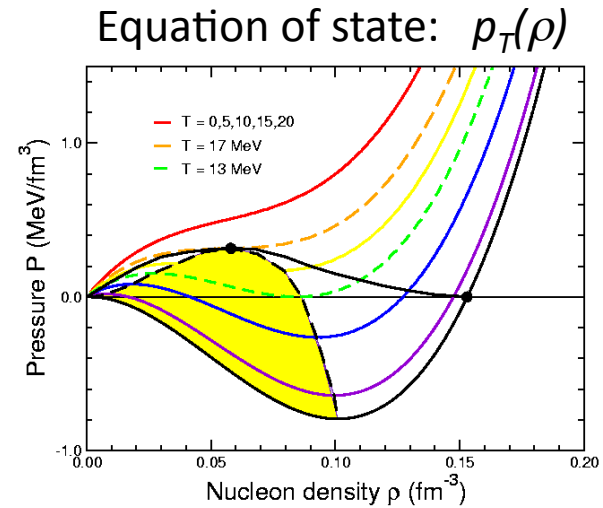
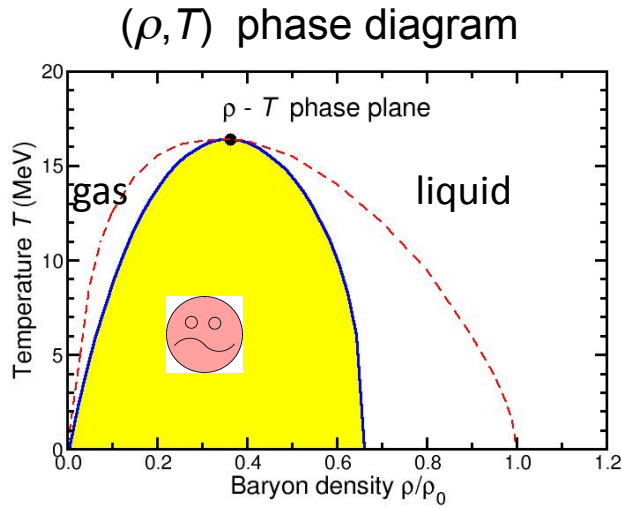
*Jørgen Randrup, LBNL*



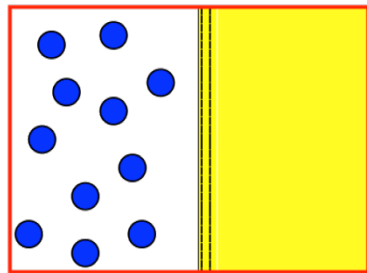
*Schematic and simplified  
phase diagram of strongly interacting matter*



# Nuclear liquid-gas phase transition



Phase coexistence

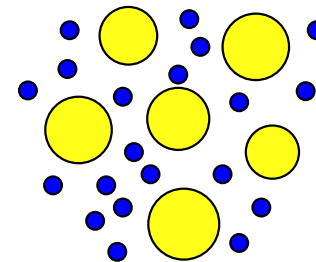


gas

liquid

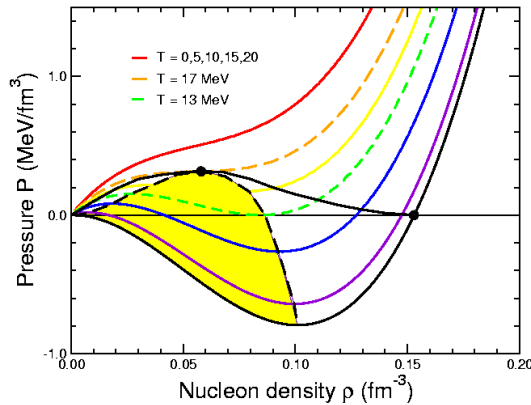
$\neq$

Phase mixture

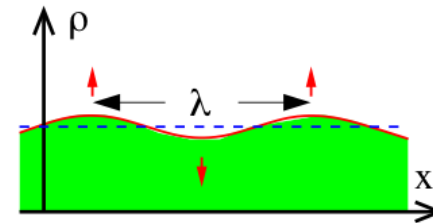


nucleons & fragments

# Nuclear spinodal fragmentation

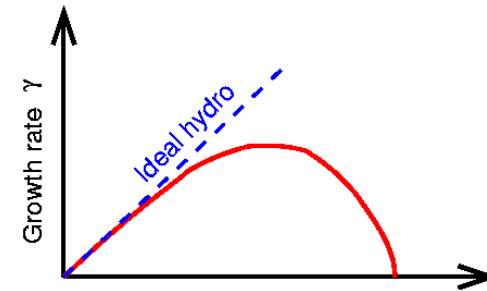


Density undulations are amplified in the spinodal region:



Long-wavelength distortions grow slowly (it takes time to relocate the matter)

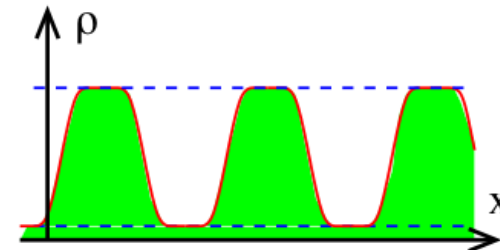
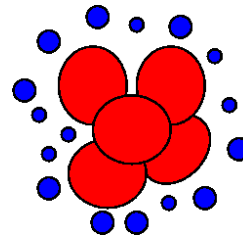
Short-wavelength distortions grow slowly (they are hardly felt due to finite range)



Wave number  $k = 2\pi/\lambda$

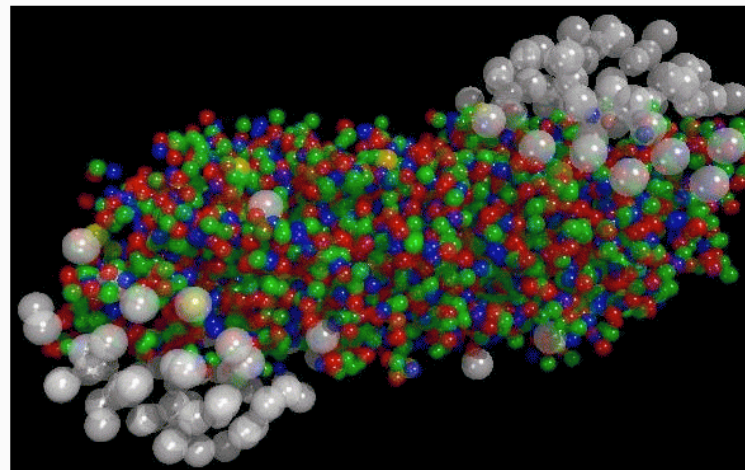
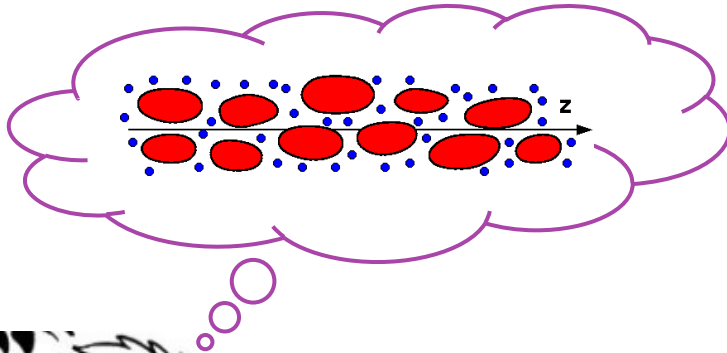
There is an *optimal length scale* that grows faster than all others:

Ph Chomaz, M Colonna, J Randrup:  
*Nuclear Spinodal Fragmentation*,  
Physics Reports 389 (2004) 263



=> *Equal-size fragments!*

*Can  
spinodal phase separation  
during the confinement  
phase transition  
happen?*



# *Spinodal phase decomposition with dissipative fluid dynamics*

Nuclear spinodal fragmentation

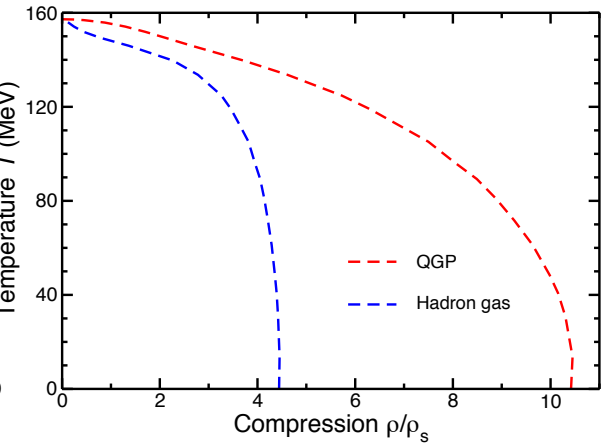
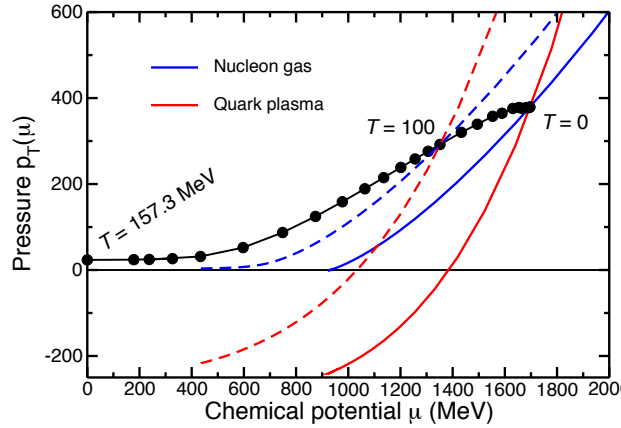
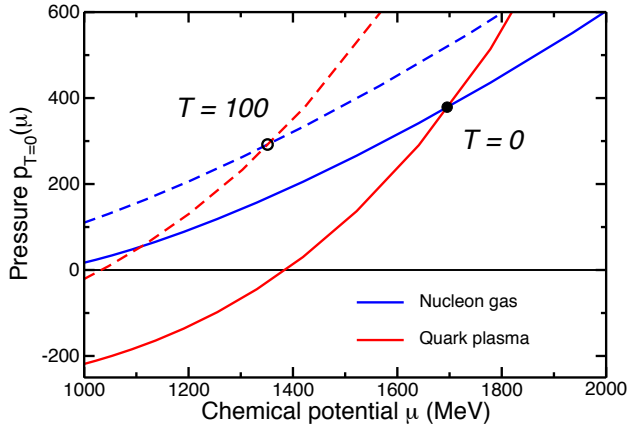
Thermodynamics

Inclusion of finite range

Dissipative fluid dynamics => dispersion relations

Dynamical phase trajectories => spinodal amplification

# Idealized equations of state: **HG** -> **QGP**



## Hadron Gas:

$$p^H = p_\pi + p_N + p_{\bar{N}} + p_w$$

$$p_\pi(T) = -g_\pi \int_{m_\pi}^{\infty} \frac{p \epsilon d\epsilon}{2\pi^2} \ln[1 - e^{-\beta\epsilon}]$$

$$p_N(T, \mu_0) = g_N \int_{m_N}^{\infty} \frac{p \epsilon d\epsilon}{2\pi^2} \ln[1 + e^{-\beta(\epsilon - \mu_0)}]$$

$$p_{\bar{N}}(T, \mu_0) = g_N \int_{m_N}^{\infty} \frac{p \epsilon d\epsilon}{2\pi^2} \ln[1 + e^{-\beta(\epsilon + \mu_0)}]$$

$$p_w(\rho) = \rho \partial_\rho w(\rho) - w(\rho)$$

$$w(\rho) = \left[ -A \left( \frac{\rho}{\rho_s} \right)^\alpha + B \left( \frac{\rho}{\rho_s} \right)^\beta \right] \rho$$

## Quark-Gluon Plasma:

$$p^Q = p_g + p_q + p_{\bar{q}} - B$$

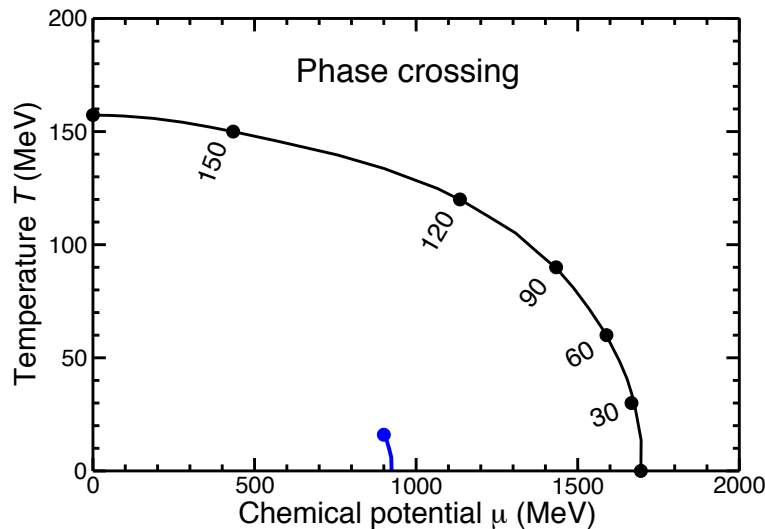
$$p_g = g_g \frac{\pi^2}{90} T^4$$

$$p_q + p_{\bar{q}} = g_q \left[ \frac{7\pi^2}{360} T^4 + \frac{1}{12} \mu_q^2 T^2 + \frac{1}{24\pi^2} \mu_q^4 \right]$$

$$\mu = \mu_0 + \partial_\rho w = 3\mu_q$$

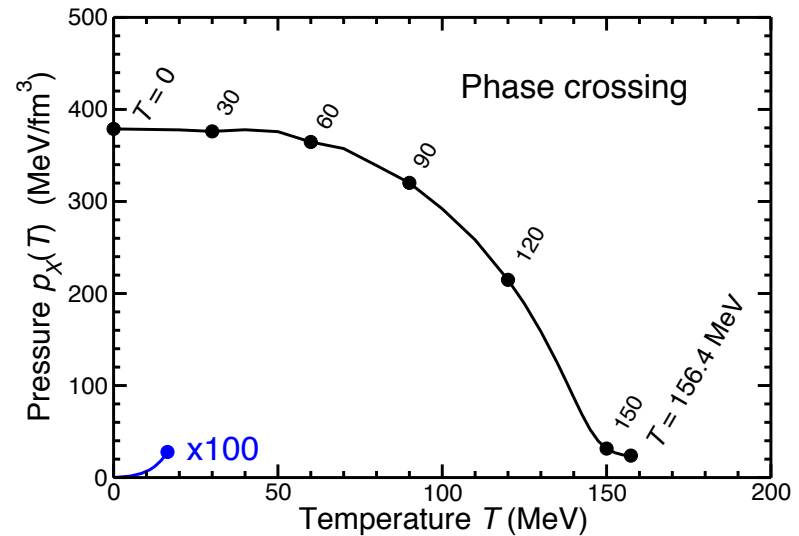
## Phase coexistence: Comparison to LG

$(\mu, T)$  phase diagram



- qualitatively similar to LG

$(T, p)$  phase diagram

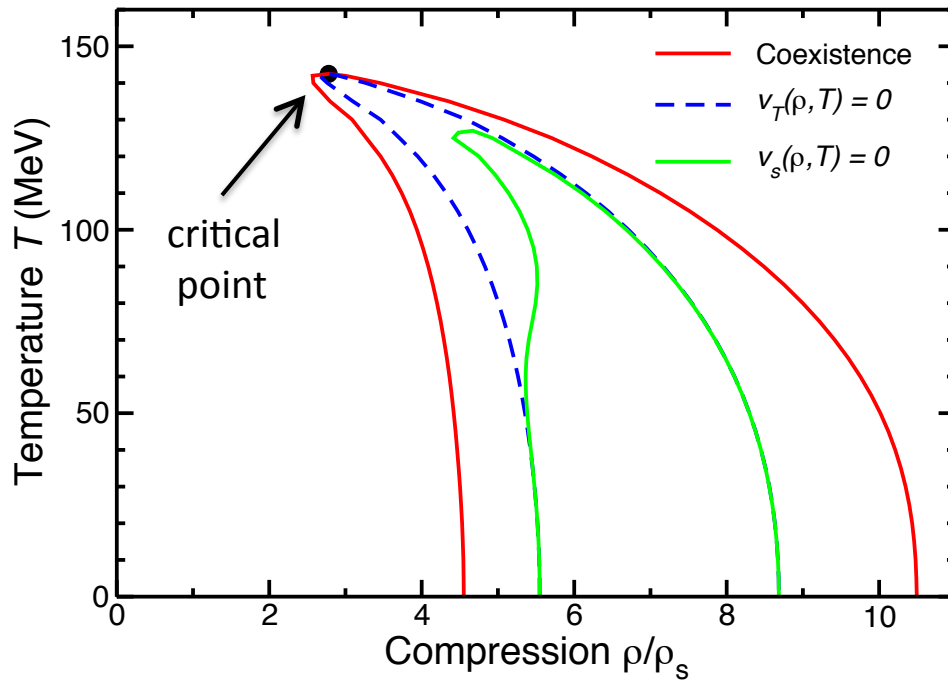


- qualitatively different from LG!

[Igor Iosilevsky, 2010]

Phase coexistence (or phase “crossing”):  
Two different media have the same  $T, \mu, p$

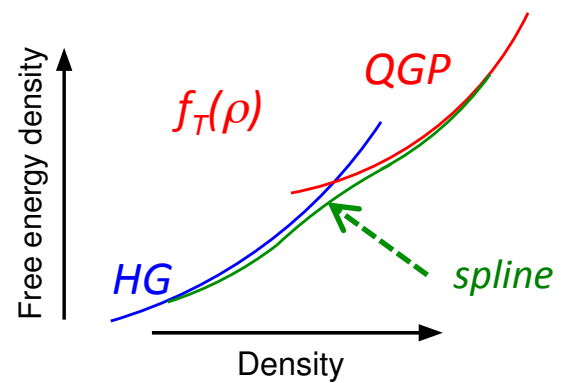
# Construct a two-phase equation of state: spline between the two idealized media



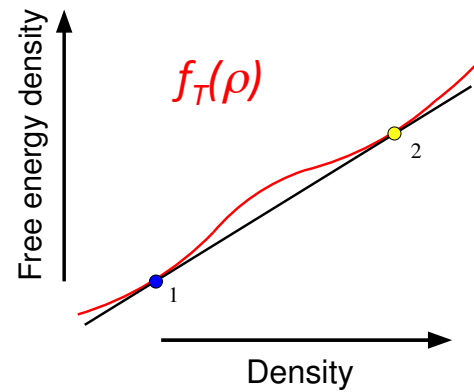
$v_T = 0$ : isothermal spinodal

$v_s = 0$ : isentropic spinodal

Spline between HG & QGP:



Phase coexistence  $\Leftrightarrow$  common tangent:



# *Spinodal phase decomposition with dissipative fluid dynamics*

Nuclear spinodal fragmentation

Thermodynamics

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Dissipative fluid dynamics => dispersion relations

Dynamical phase trajectories => spinodal amplification

## Equation of state: Finite range

Free energy density for uniform matter:  $f_0(\rho, T)$

But we need to treat non-uniform systems:  $\tilde{\rho}(\mathbf{r}), \tilde{T}(\mathbf{r})$

The local density approximation:  $f[\tilde{\rho}(\cdot), \tilde{T}(\cdot)](\mathbf{r}) \doteq f_0(\tilde{\rho}(\mathbf{r}), \tilde{T}(\mathbf{r}))$

... implies:

$$F(\text{[uniform block]}) = F(\text{[discrete bars]})$$

No good!  $\Rightarrow$  Finite range *must* be taken into account

## Non-uniform density $\tilde{\rho}(\mathbf{r})$

<i>gradient contribution</i>	$\tilde{w}(\mathbf{r}) \equiv w_0(\tilde{\rho}(\mathbf{r})) + \frac{1}{2}C(\nabla\tilde{\rho}(\mathbf{r}))^2$
------------------------------	---

local entropy density:  $\tilde{\sigma}(\mathbf{r}) \equiv \sigma(\tilde{\epsilon}(\mathbf{r}), \tilde{\rho}(\mathbf{r}))$

$\Rightarrow$  total entropy:  $S[\tilde{\epsilon}(\mathbf{r}), \tilde{\rho}(\mathbf{r})] \equiv \int \tilde{\sigma}(\mathbf{r}) d\mathbf{r}$

$$\Rightarrow \begin{cases} \tilde{\beta}(\mathbf{r}) \equiv \frac{\delta S}{\delta \tilde{\epsilon}(\mathbf{r})} \Rightarrow \tilde{T}(\mathbf{r}) \\ \tilde{\alpha}(\mathbf{r}) \equiv \frac{\delta S}{\delta \tilde{\rho}(\mathbf{r})} \Rightarrow \tilde{\mu}(\mathbf{r}) \end{cases}$$

$\Rightarrow$  local pressure  $\tilde{p}(\mathbf{r})$  & local enthalpy density  $\tilde{h}(\mathbf{r})$  & ...

Free energy density:  $\tilde{f}_T(\mathbf{r}) = f_T(\tilde{\rho}(\mathbf{r})) + \frac{1}{2}C(\nabla\tilde{\rho}(\mathbf{r}))^2$

J. Randrup, Phys. Rev. C 79, 054911 (2009)

H. Heiselberg *et al.*, Phys. Rev. Lett. 70, 1355 (1993)

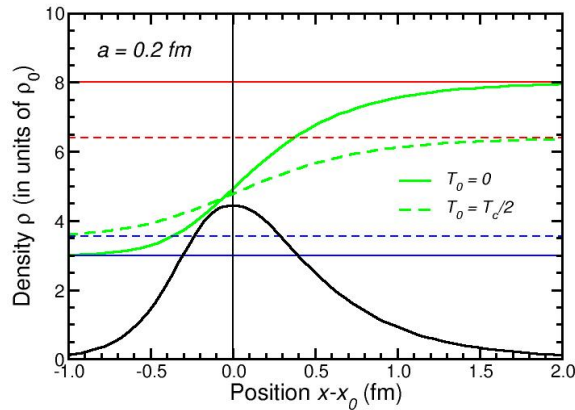
# Interface equilibrium



Global equilibrium requires constant  $T, \mu, \rho$ :

$$0 \doteq \delta S - \beta_0 \delta E - \alpha_0 \delta N = \int dx \left\{ [\tilde{\beta}(x) - \beta_0] \delta \tilde{\epsilon}(x) + [\tilde{\alpha}(x) - \alpha_0] \delta \tilde{\rho}(x) \right\}$$

$$\Rightarrow \tilde{\beta}(x) = \beta_0 \ \& \ \tilde{\alpha}(x) = \alpha_0 \ \Rightarrow \ \tilde{p}(x) = p_0$$



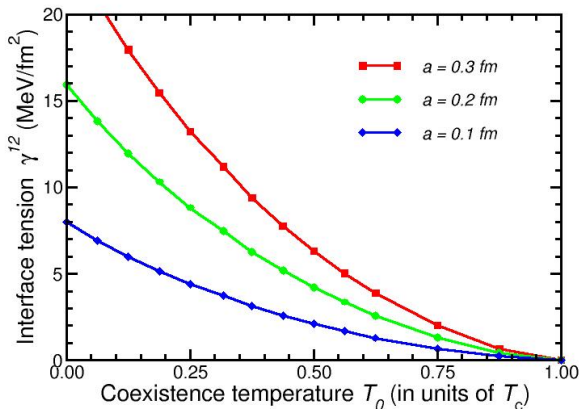
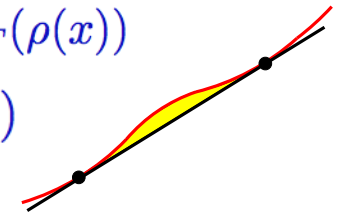
$$\tilde{f}_T(\mathbf{r}) = f_T(\tilde{\rho}(\mathbf{r})) + \frac{1}{2} C (\nabla \tilde{\rho}(\mathbf{r}))^2 \quad \Rightarrow$$

The interface density profile is determined by

$$C \partial_x^2 \rho(x) \doteq \mu_T(\rho(x)) - \mu_0 = \partial_\rho \Delta f_T(\rho(x))$$

where  $\Delta f_T(\rho) \equiv f_T(\rho) - f_T^M(\rho)$

$$f_T^M(\rho) \equiv f_T(\rho_i) + \mu_0(\rho - \rho_i) \leq f_T(\rho)$$



The interface tension is given by

$$\gamma_T^{12} = \int_{\rho_1}^{\rho_2} d\rho [2C \Delta f_T(\rho)]^{\frac{1}{2}}$$

[The density profile  $\rho(x)$  is not needed!]

J. Randrup, Phys. Rev. C 79, 054911 (2009)

*Dynamical effect of the gradient term:  
The local pressure is modified*

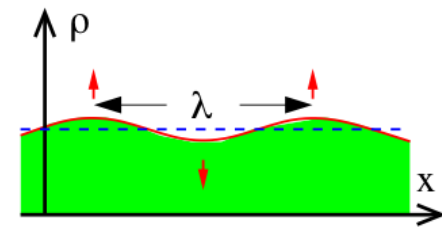
Small deviations from uniformity:

$$\tilde{p}(\mathbf{r}) \approx p_0(\tilde{\varepsilon}(\mathbf{r}), \tilde{\rho}(\mathbf{r})) - C\rho_0\nabla^2\tilde{\rho}(\mathbf{r})$$

Small harmonic density undulations:

$$\tilde{\rho}(\mathbf{r}, t) = \rho_0 + \delta\rho(x, t) \doteq \rho_0 + \rho_k e^{ikx - i\omega t}$$

$$p_k \rightarrow p_k + C\rho_0 k^2 \rho_k$$



# *Spinodal phase decomposition with dissipative fluid dynamics*

Nuclear spinodal fragmentation

Thermodynamics

Inclusion of finite range

Dissipative fluid dynamics => dispersion relations

Dynamical phase trajectories => spinodal amplification

# Transport model: Dissipative fluid dynamics

Energy-momentum tensor:

$$T^{00} \approx \varepsilon \quad \& \quad T^{0i} \approx (\varepsilon + p)v^i + q^i \quad \& \quad \text{A Muronga, PRC 76, 014909 (2007)}$$

$$T^{ij} \approx \delta_{ij}p - \eta[\partial_i v^j + \partial_j v^i - \frac{2}{3}\delta_{ij}\partial^k v^k] - \zeta\delta_{ij}\partial^k v^k \quad |\rho_k| \ll \rho_0 \Rightarrow |v| \ll 1$$

$$\nabla \cdot \mathbf{T} \approx \nabla p - \eta\Delta\mathbf{v} - [\frac{1}{3}\eta + \zeta]\nabla(\nabla \cdot \mathbf{v}) \asymp \partial_x p - [\frac{4}{3}\eta + \zeta]\partial_x^2 v \quad \text{Eckart frame}$$

Equations of motion:

$$\left\{ \begin{array}{l} C : \partial_t \rho \doteq -\rho_0 \nabla \cdot \mathbf{v} \Rightarrow \omega \rho_k \doteq \rho_0 k v_k \quad \text{charge} \\ M : h_0 \partial_t \mathbf{v} \doteq -\nabla[p - \zeta \nabla \cdot \mathbf{v}] - \nabla \cdot \boldsymbol{\pi} - \partial_t \mathbf{q} \quad \text{momentum} \\ E : \partial_t \varepsilon \doteq -h_0 \nabla \cdot \mathbf{v} - \nabla \cdot \mathbf{q} \quad \text{energy} \end{array} \right.$$

Sound equation:

$$\partial_t E - \nabla \cdot \mathbf{M} : h_0 \partial_t^2 \varepsilon \doteq \Delta[p - \zeta \nabla \cdot \mathbf{v}] + \nabla \cdot (\nabla \cdot \boldsymbol{\pi})$$

$$\omega^2 \varepsilon_k \doteq k^2 p_k - i[\frac{4}{3}\eta + \zeta] \frac{\omega}{\rho_0} k^2 \rho_k \quad \xi \equiv \frac{4}{3}\eta + \zeta$$

Heat flow:

$$\mathbf{q} \approx -\kappa[\nabla T + T_0 \partial_t \mathbf{v}] : q_k = -i\kappa[kT_k - \frac{T_0}{\rho_0} \frac{\omega^2}{k} \rho_k] \quad T_k \approx \frac{1}{1 + i\kappa k^2 / \omega c_v} \frac{T_0}{\rho_0} \left( \frac{\partial p}{\partial \varepsilon} \right)_\rho \rho_k$$

Equation of state:

$$p_T(\rho) \Rightarrow p_k = \left( \frac{\partial p}{\partial \varepsilon} \right)_\rho c_v T_k + \frac{h_0}{p_0} v_T^2 \rho_k$$

Dispersion equation:

$$\omega^2 \doteq v_T^2 k^2 + C \frac{\rho_0^2}{h_0} k^4 - i\xi \frac{\omega}{h_0} k^2 + \frac{v_s^2 - v_T^2}{1 + i\kappa k^2 / \omega c_v} k^2$$

Heiselberg, Pethick, Ravenhall,  
Ann. Phys. 233, 37 (1993)

# Dispersion relations

$$\omega_k = \epsilon_k + i\gamma_k$$

Ideal fluid dynamics:

$$\omega^2 \doteq v_s^2 k^2$$

+ gradient term:

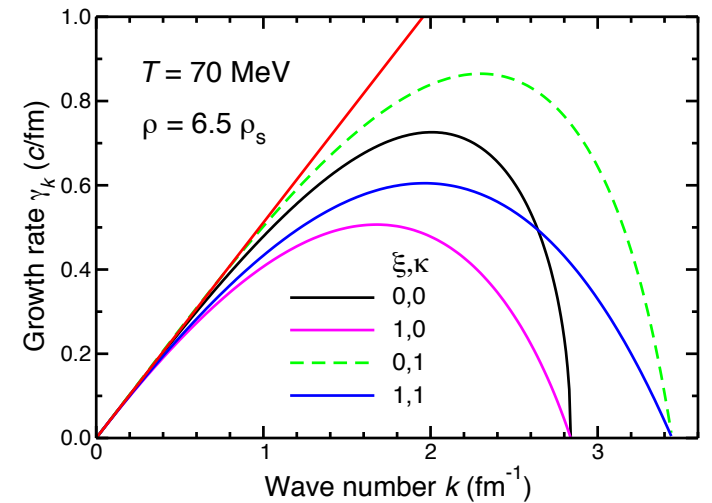
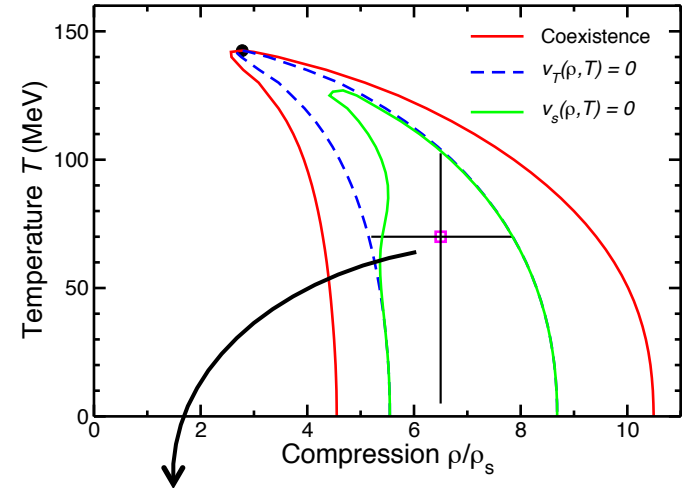
$$\omega^2 \doteq v_s^2 k^2 + C \frac{\rho_0^2}{h_0} k^4$$

+ shear & bulk viscosity:

$$\omega^2 \doteq v_s^2 k^2 + C \frac{\rho_0^2}{h_0} k^4 - i\xi \frac{\omega}{h_0} k^2$$

+ heat conduction:

$$\omega^2 \doteq v_T^2 k^2 + C \frac{\rho_0^2}{h_0} k^4 - i\xi \frac{\omega}{h_0} k^2 + \frac{v_s^2 - v_T^2}{1 + i\kappa k^2 / \omega c_v} k^2$$



$$\xi \equiv \frac{4}{3}\eta + \zeta$$

## Transport coefficients

$$\eta_0 \geq 1$$

$$\kappa_0 \geq 1$$

1) Bulk viscosity  $\zeta$ : Ignore  $\zeta \ll \eta \Rightarrow \xi \equiv \frac{4}{3}\eta + \zeta \approx \frac{4}{3}\eta$

2) Shear viscosity  $\eta$ :

$$\rho = 0 : h \equiv p + \varepsilon = T\sigma$$

$$\rho = 0 : \eta \geq \frac{\hbar}{4\pi} \sigma = \frac{\hbar}{4\pi} \frac{h}{T}$$

$$\rho > 0, T \ll mc^2 : h \asymp mc^2 n \gg T\sigma$$

$$\rho = 0 : n \sim T^3 \Rightarrow \frac{\hbar c}{T} = 4\pi c_0 d \quad d \equiv n^{1/3}$$

$$\eta(\rho, T) = \eta_0 \frac{c_0}{c} d(\rho, T) h(\rho, T)$$

$$\lambda_{\text{visc}} \equiv \frac{1}{c} \frac{\xi(\rho, T)}{h(\rho, T)/c^2} \approx \frac{4}{3} \eta_0 c_0 d(\rho, T)$$

3) Heat conductivity  $\kappa$ :

$$\eta \approx \frac{1}{3} n \bar{p} \ell$$

$$\kappa \approx \frac{1}{3} \bar{v} \ell c_v$$

$$\frac{\kappa}{\eta} \approx \frac{c_v}{h/c^2}$$

$$\bar{p} = m\bar{v}$$

$$h \asymp mc^2 n$$

$$c_v \equiv \partial_T \varepsilon_T(\rho)$$

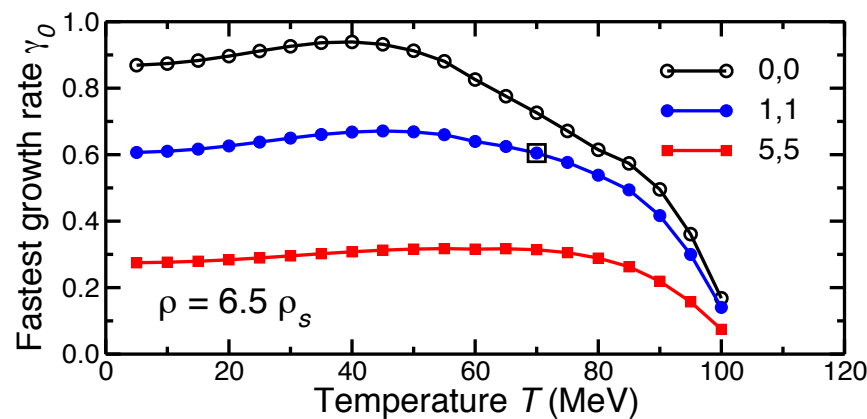
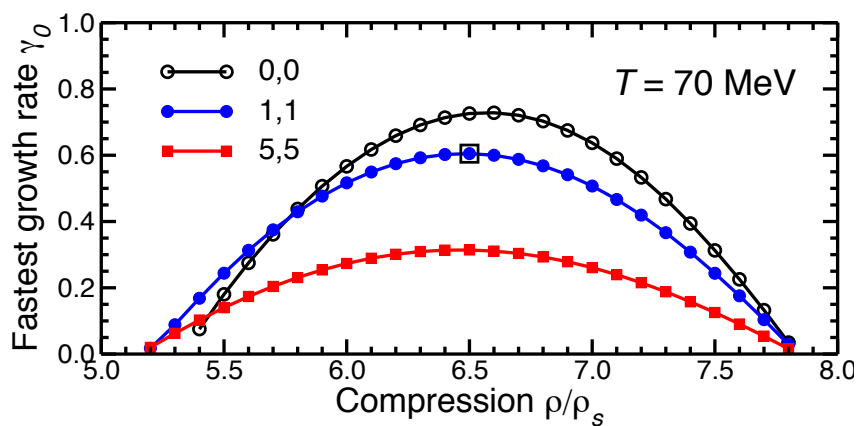
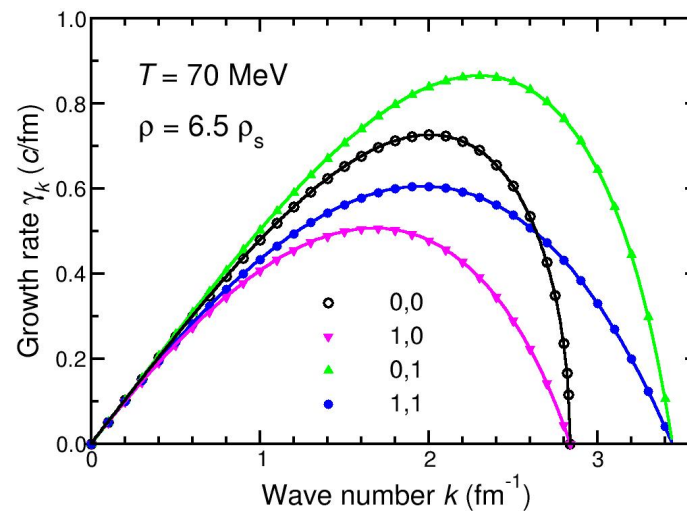
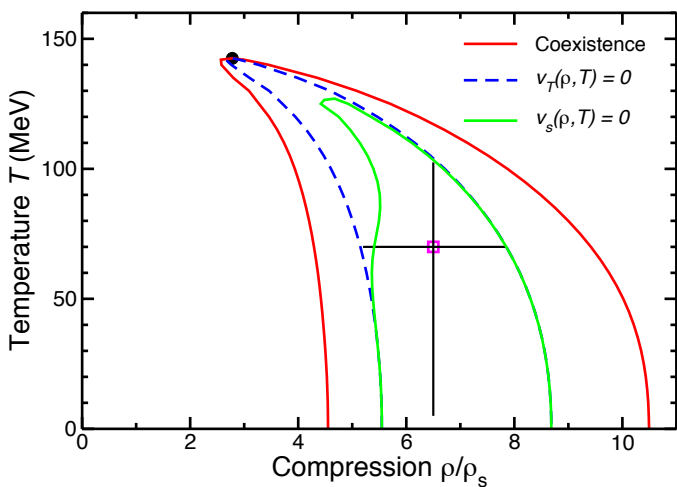
$$c_v \asymp \frac{3}{2} n$$

$$\kappa(\rho, T) = \kappa_0 c_0 c d(\rho, T) c_v(\rho, T)$$

$$\lambda_{\text{heat}} \equiv \frac{1}{c} \frac{\kappa(\rho, T)}{c_v(\rho, T)} = \kappa_0 c_0 d(\rho, T)$$

$$c_0 = \frac{1}{4\pi} \left[ (g_g + \frac{3}{3} g_q) \frac{\zeta(3)}{\pi^2} \right]^{\frac{1}{3}} \approx 0.12779$$

# Spinodal growth rates



# *Spinodal phase decomposition with dissipative fluid dynamics*

Nuclear spinodal fragmentation

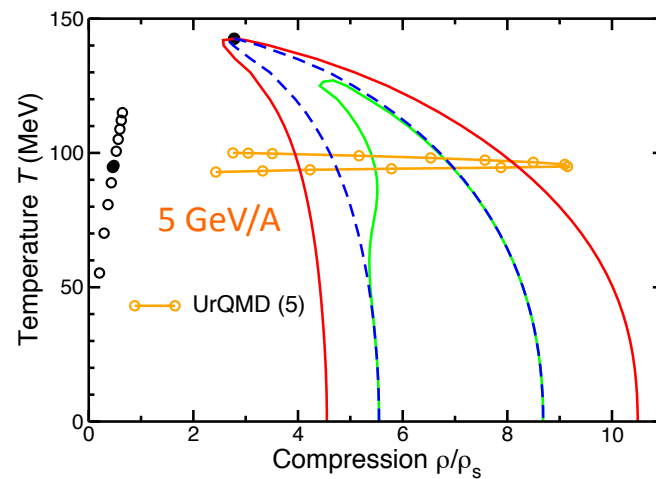
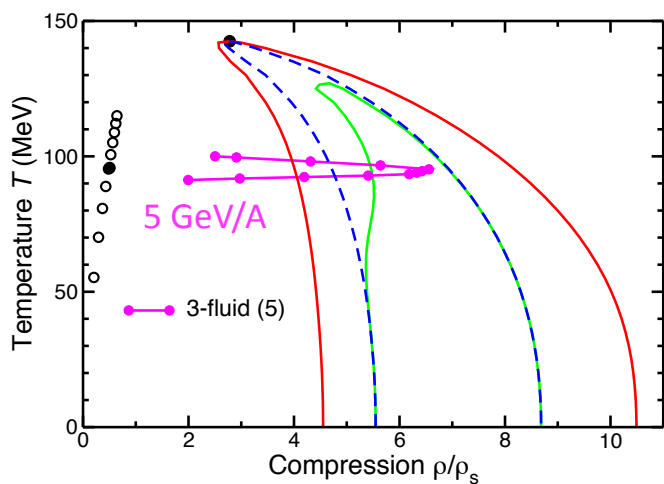
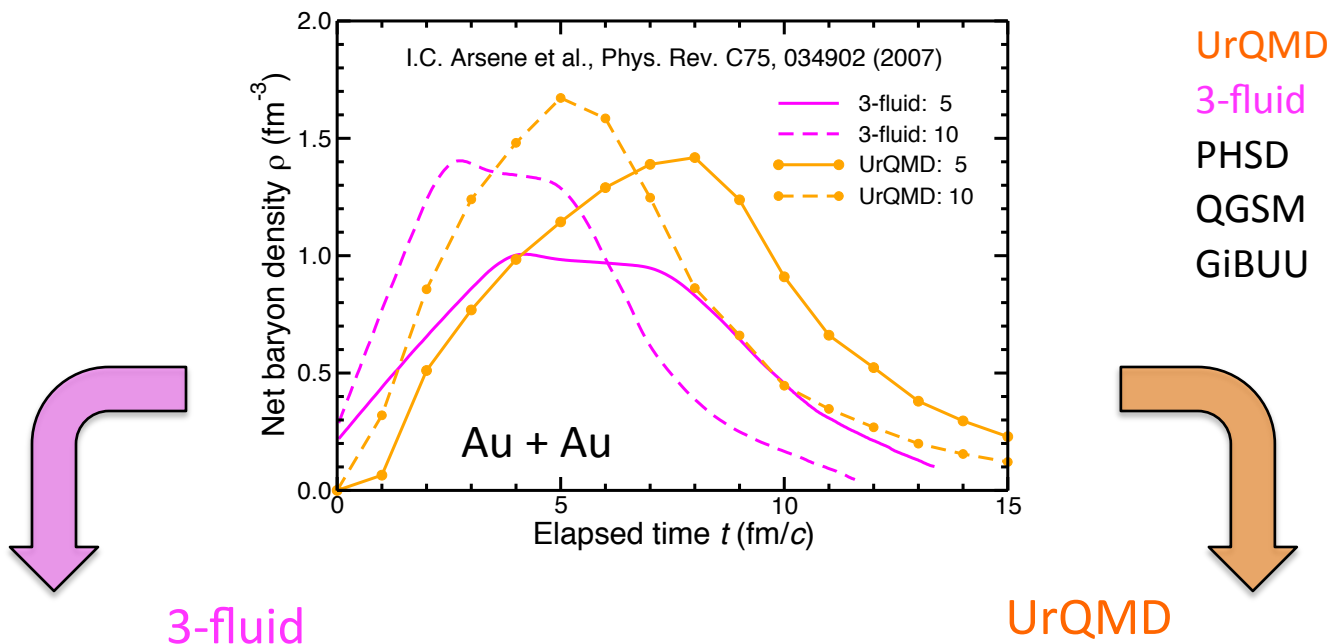
Thermodynamics

Inclusion of finite range

Dissipative fluid dynamics => dispersion relations

Dynamical phase trajectories => spinodal amplification

# Dynamical phase trajectories



# Dynamics of collective modes in many-body systems

Amplitude evolution:

$$\frac{d}{dt}A_\nu(t) = -i\omega_\nu A_\nu(t) + B_\nu(t)$$

$\omega_\nu = \epsilon_\nu + i\gamma_\nu$

Correlation function:

$$\sigma_{\nu\mu}(t_1, t_2) \equiv \langle A_\nu(t_1) A_\mu(t_2)^* \rangle$$

Markovian noise:

$$\langle B_\nu(t) B_\mu(t')^* \rangle = 2\mathcal{D}_{\nu\mu} \delta(t - t')$$

Evolution:

$$\frac{d}{dt}\sigma_{\nu\mu}(t) = 2\mathcal{D}_{\nu\mu} - i(\omega_\nu - \omega_\mu^*)\sigma_{\nu\mu}$$

*seed*                      *feedback*

Variance of a single mode:

$$\frac{d}{dt}\sigma_\nu^2 = 2\mathcal{D}_\nu + 2\gamma_\nu\sigma_\nu^2$$

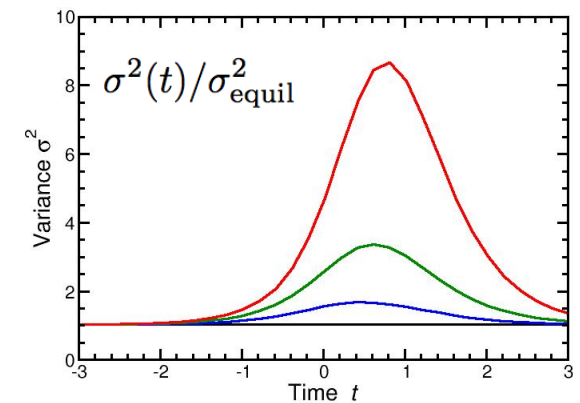
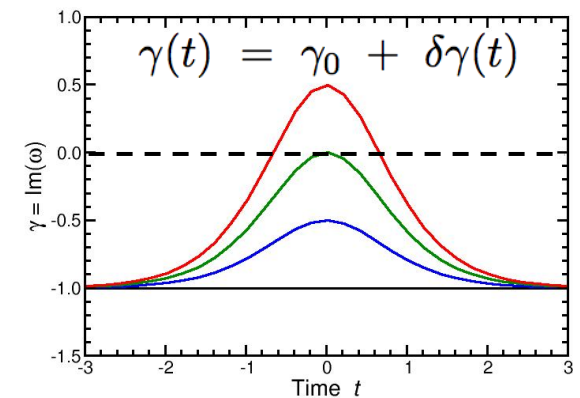
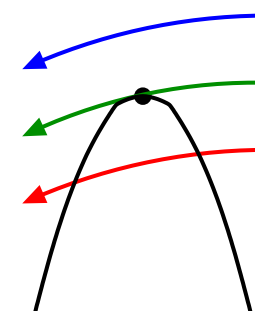
*Lalime Equation*

*Amplification coefficient:*

$$\Gamma_\nu(t) \equiv \int_0^t \gamma_\nu(t') dt'$$

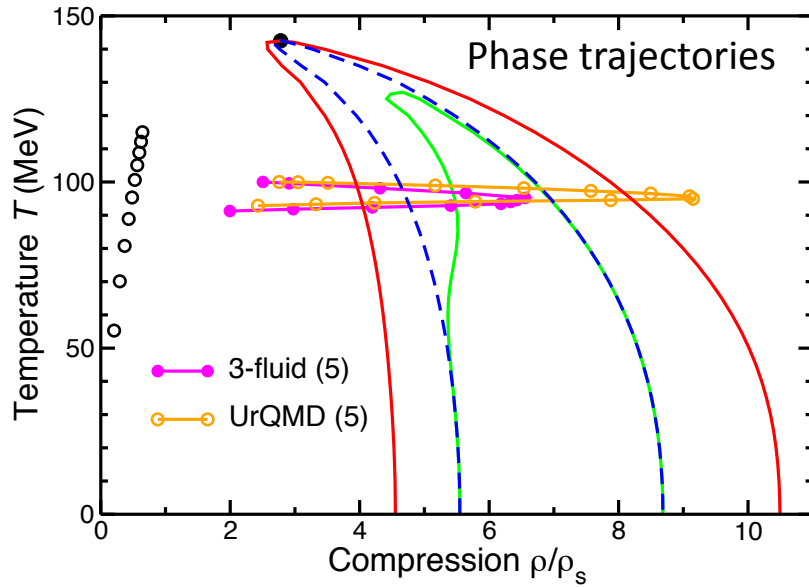
$$\Rightarrow \sigma_\nu^2(t) = \left[ 2\mathcal{D}_\nu \int_0^t e^{-2\Gamma_\nu(t')} dt' + \sigma_0^2 \right] e^{2\Gamma_\nu(t)}$$

$$\gamma_\nu < 0 : \sigma_\nu^2(t) \rightarrow -\mathcal{D}_\nu / \gamma_\nu$$



Colonna, Chomaz, Randrup, NPA 567 (1994) 637

# Spinodal amplification

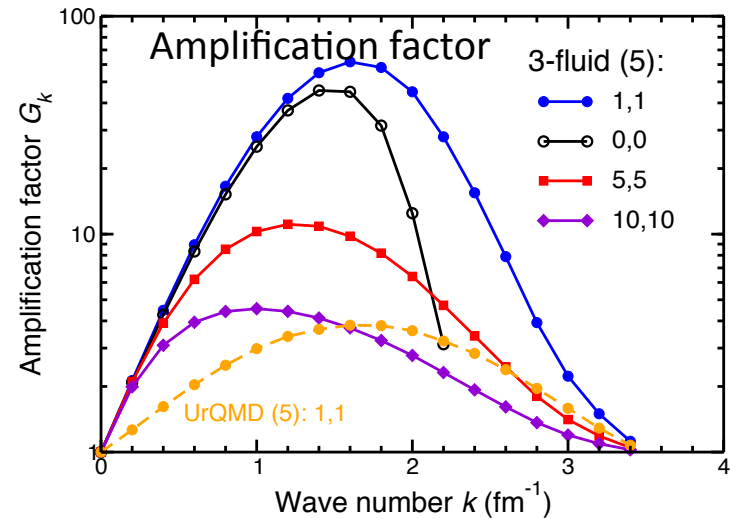
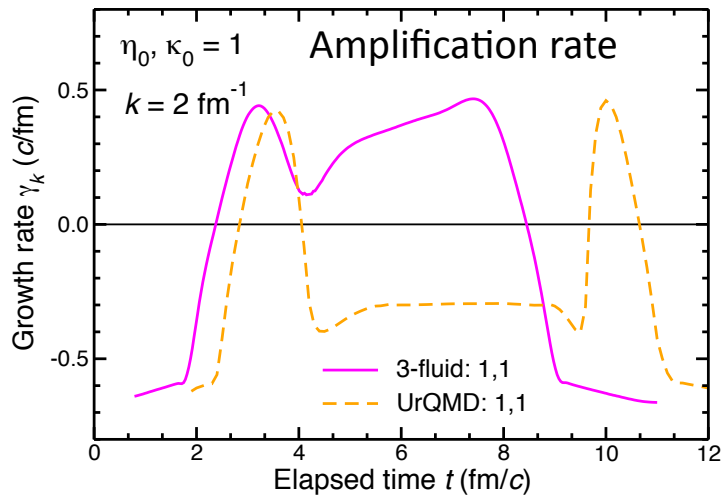


Amplification coefficient:

$$\Gamma_\nu(t) \equiv \int_0^t \gamma_\nu(t') dt'$$

Amplification factor:

$$G_k = e^{2\Gamma_k}$$



Can spinodal phase separation occur during confinement?

# Conclusions

Spinodal decomposition may occur at  $E_{\text{lab}} \approx 5\text{-}10$  GeV/A (FAIR & NICA):

- 1) Exists an optimal  $E_{\text{lab}}$  range
- 2) Significant amplification

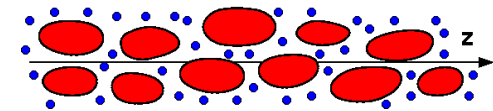
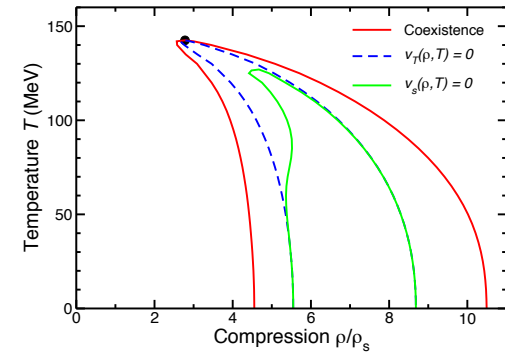
- but:

*Full dynamical calculations are needed!*

Suggestion: Use fluid dynamics with finite range

Need: Transport coefficients  $\eta(\rho, T)$ ,  $\zeta(\rho, T)$ ,  $\kappa(\rho, T)$  in addition to the Equation of State  $p(\rho, T)$

If it occurs: Useful signal of phase transition



## MODEL REQUIREMENTS:

EoS with phase transition

Confined & deconfined phases

Phase coexistence, incl interface\*

Spinodal modes\*

Dynamics with instabilities\*

\* Requires finite range