

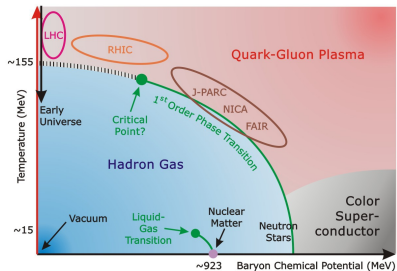
Dynamic Critical Behaviour and Spectral Functions of ϕ^4 Theory (A06)

D. Schweitzer¹, S. Schlichting², L. v. Smekal¹

¹Institut für Theoretische Physik, Justus-Liebig-Universität Gießen

²Fakultät für Physik, Universität Bielefeld

- Long-term objective: Find the critical point of QCD
 - Experiments: heavy-ion collisions
 - Naturally very dynamic
 - Equilibrium not guaranteed
- ⇒ Need to know how critical point changes dynamics
- Accessible dynamic quantity: spectral functions



- *Continuous* phase transition
 - Infinite correlation length ξ
 - Scale invariance
 - ⇒ Universality
- Power laws in observables ($\langle \phi \rangle, \chi, \xi, \dots$):

$$\xi \sim |\tau|^{-\nu}$$
$$\langle O(T) \rangle \sim |\tau|^\sigma, \quad \tau := \frac{T - T_c}{T_c}$$

- Examples: CO_2 , ferromagnets, QCD, $\phi^4 \dots$

Fully described by classical physics!



Figure: "Critical opalescence of ethane" by Dr. Sven Horstmann, used under CC BY-SA 3.0/cropped from original.

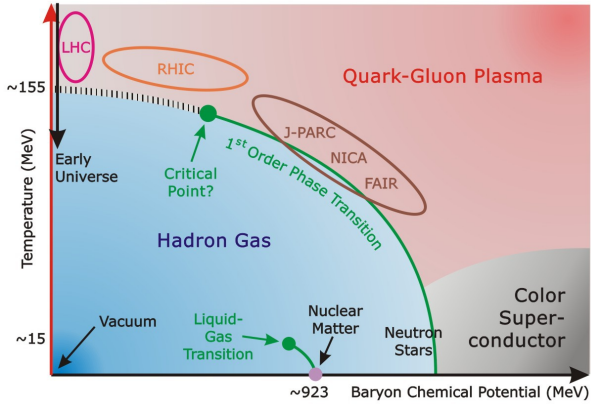


Figure: Semi-quantitative phase diagram of QCD

$$\mathcal{L}(\phi) = \frac{1}{2} (\partial^\mu \phi \partial_\mu \phi - m^2 \phi^2) - \frac{\lambda}{4!} \phi^4 - J\phi$$

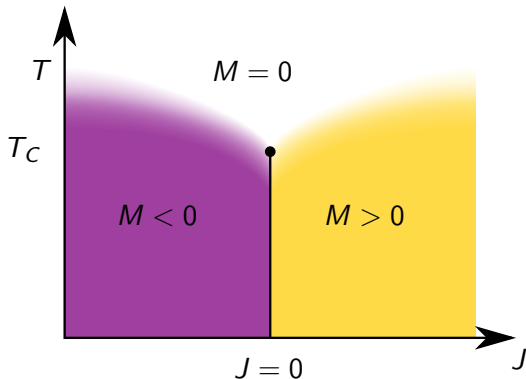


Figure: Phase diagram of a ϕ^4 theory. The line between opposing magnetization phases ends in a critical point at the curie temperature.

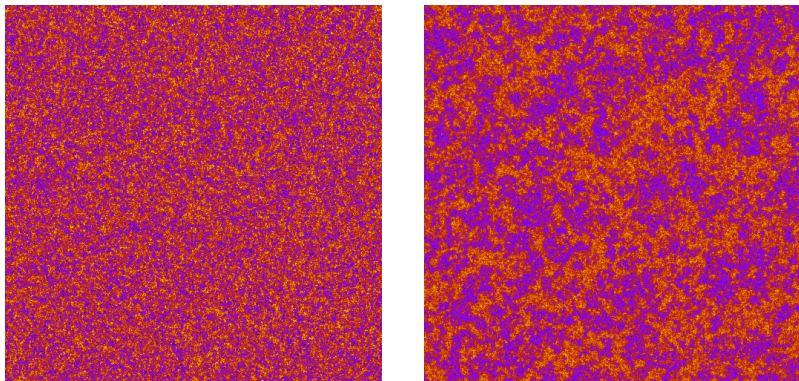


Figure: Configurations in the symmetric phase

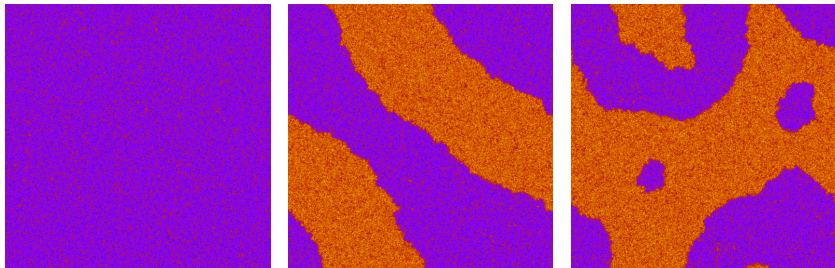


Figure: Configurations on line of phase coexistence

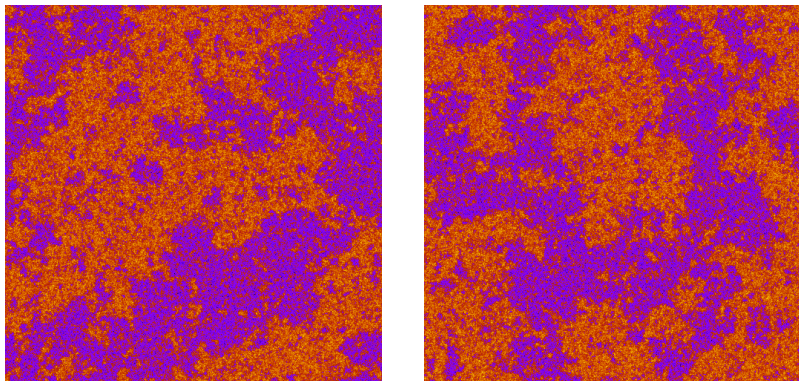


Figure: Configurations close to the critical point

- Time evolution of expectation values:

$$\langle \hat{O}(t) \rangle = Z^{-1} \text{tr} \left(\hat{\rho}_0 e^{-i\hat{H}t} \hat{O} e^{i\hat{H}t} \right) \quad (1)$$

- Suzuki-Trotter decomposition \Rightarrow *Schwinger-Keldysh* technique

$$\langle \hat{O}(t) \rangle = Z^{-1} \int \mathcal{D}\phi_+(t) \mathcal{D}\phi_-(t) \rho_0[\phi_+(0), \phi_-(0)] \exp \left[-i \int_0^t dt_1 \hat{H}(\phi_+(t_1)) \right] \mathcal{O}[\phi_+(t), \phi_-(t)] \quad (2)$$

$$\exp \left[+i \int_t^0 dt_2 \hat{H}(\phi_-(t_2)) \right] = \int \mathcal{D}\phi_+ \mathcal{D}\phi_- \rho_0 e^{-i \int_0^t dt' \frac{\dot{\phi}_+^2 - \dot{\phi}_-^2}{2} - (V(\phi_+) - V(\phi_-))} \mathcal{O}(\phi_+, \phi_-) \quad (3)$$

- Change of variables:
 - “classical” coordinate $\phi = \frac{\phi_+ + \phi_-}{2}$
 - “quantum fluctuation” $\tilde{\phi} = \phi_+ - \phi_-$
- Expand (3) to first order in $\tilde{\phi}$:

$$\langle \hat{O}(t) \rangle \approx \int \mathcal{D}\phi \mathcal{D}\tilde{\phi} e^{-i \int_0^t dt' \tilde{\phi}(t') (-\ddot{\phi}(t') - \partial_\phi V(\phi))} \rho_0(\phi, \dot{\phi}) O(\phi, \dot{\phi})$$

- \tilde{x} -integration enforces classical equations of motion: $\ddot{\phi} + \partial_\phi V(\phi) = 0$

$$H(\phi) = \frac{1}{2} \dot{\phi}^2 + V(\phi)$$

$$V(\phi) = \frac{1}{2} ((\nabla\phi)^2 + m^2\phi^2) + \frac{\lambda}{4!} \phi^4 - J\phi$$

$$\phi(t + \Delta t) - \phi(t) = \Delta t \cdot \dot{\phi}$$

$$\dot{\phi}(t + \Delta t) - \dot{\phi}(t) = -\Delta t \cdot \partial_{\phi} H$$

- Discretize $\phi(x_i), \dot{\phi}(x_i)$ on square/cubic lattice
 - Solve equations of motion at every lattice site
 - Hamiltonian dynamics: energy conserved \rightarrow microcanonical ensembles
 - Average over thermal initial conditions
- \Rightarrow Calculate real-time observables as functions of $\phi(t), \dot{\phi}(t)$

$$H(\phi) = \frac{1}{2} \dot{\phi}^2 + V(\phi)$$

$$V(\phi) = \frac{1}{2} ((\nabla\phi)^2 + m^2\phi^2) + \frac{\lambda}{4!} \phi^4 - J\phi$$

$$\phi(t + \Delta t) - \phi(t) = \Delta t \cdot \dot{\phi}$$

$$\dot{\phi}(t + \Delta t) - \dot{\phi}(t) = -\Delta t \cdot \partial_{\phi} H - \Delta t \left(\gamma \dot{\phi} + \sqrt{2\gamma T} \eta(t) \right)$$

- Discretize $\phi(x_i), \dot{\phi}(x_i)$ on square/cubic lattice
 - Solve equations of motion at every lattice site
 - **Langevin dynamics: coupling to heat bath \rightarrow canonical ensembles**
 - Average over thermal initial conditions
- \Rightarrow Calculate real-time observables as functions of $\phi(t), \dot{\phi}(t)$

WARNING

The following footage may potentially trigger seizures for people with photosensitive epilepsy. Viewer discretion is advised.

- Passing of critical point changes dynamics visibly
 - Known phenomenon: Critical Slowing-Down (of MC algorithms)
- ⇒ Look at multi-time correlation functions

- Spectral function defined via decomposition of Green's function:

$$\begin{aligned}\rho(t, t', x, x') &\equiv i \langle [\phi(t, x) \phi(t', x')] \rangle \\ F(t, t', x, x') &\equiv \frac{1}{2} \langle \{ \phi(t, x) \phi(t', x') \} \rangle - \langle \phi \rangle^2\end{aligned}$$

- Use fluctuation-dissipation theorem to get ρ from F

$$F(\omega, \vec{p}) = -i \left(\frac{1}{2} + n_T(\omega) \right) \rho(\omega, \vec{p}) \quad (4)$$

- Use classical limit of BE distribution $n_T(\omega) \approx \frac{T}{\omega}$ for large $\frac{T}{\omega}$:

$$F(\omega, \vec{p}, T) = -i \frac{T}{\omega} \rho(\omega, \vec{p}, T) \quad (5)$$

$$\Rightarrow \rho(t, \vec{p}, T) = -\frac{1}{T} \frac{\partial}{\partial t} F(t, \vec{p}, T) \quad (6)$$

$$= -\frac{1}{T} \left\langle \dot{\phi}(t - t', p) \phi(0, -p) \right\rangle \quad (7)$$

Now calculate ρ from classical fields $\phi, \dot{\phi}$!

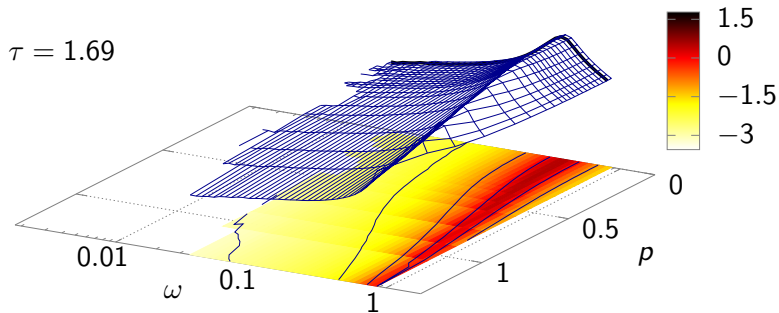


Figure: Spectral functions at different temperatures

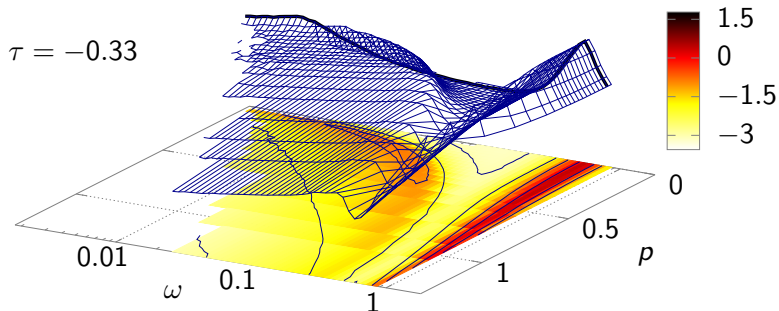


Figure: Spectral functions at different temperatures

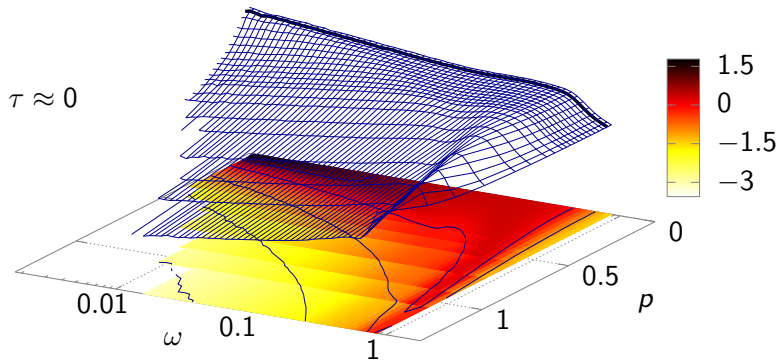


Figure: Spectral functions at different temperatures

- Time scale diverges with length scale: $\xi_t \sim \xi^z \sim \tau^{-\nu z}$
- Time/frequency variables only enter observables in rescaled form:

$$\tau \equiv \frac{T - T_c}{T_c}$$

$$\rho(\omega, p, \tau) = s^{(2-\eta)} \rho\left(s^z \omega, sp, s^{\frac{1}{\nu}} \tau\right) \quad (8)$$

- Choose $s = \omega^{-1/z}$

$$\rho(\omega, p, \tau) = \omega^{-(2-\eta)/z} \tilde{\rho}\left(p/\omega^{1/z}, \tau/\omega^{1/\nu z}\right) \quad (9)$$

- ⇒ Universal scaling function $\tilde{\rho}$
- ⇒ IR power law at $p = 0, \tau \rightarrow 0$:

$$\rho(\omega, p, \tau) = \omega^{-(2-\eta)/z} \cdot \text{const.} \quad (10)$$

- Fourier transform (9) to time domain:

$$\rho(t, p, \tau) = t^{-\frac{2-\eta}{z}-1} \tilde{\rho}_t \left(p t^{1/z}, \tau / t^{1/\nu z} \right) \quad (11)$$

- Ansatz for $\tilde{\rho}_t$ at vanishing spatial momentum:

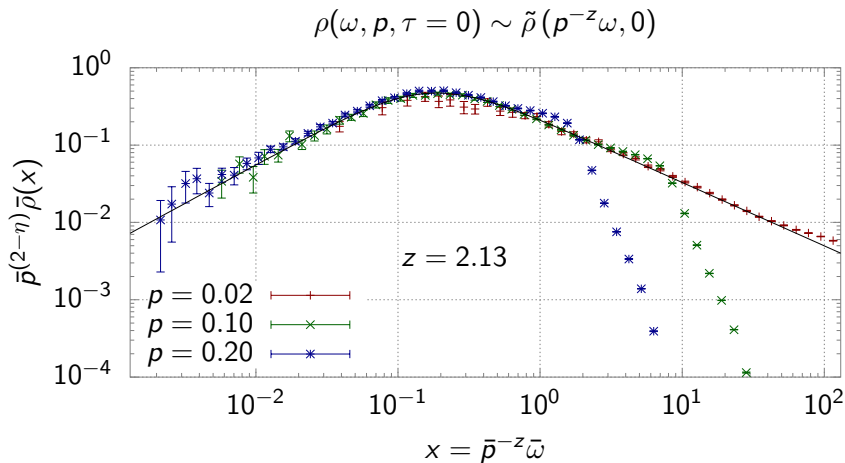
$$\tilde{\rho}_t \left(0, \tau / t^{1/\nu z} \right) \propto \exp \left(-\frac{t}{\xi_t} \right) \quad (12)$$

Additional quantities:

Dynamic critical exponent z , correlation time ξ_t

- Time/frequency scaling exponent z
⇒ dynamic universality classes
- Additional influences on z :
 - Conservation laws
 - Poisson brackets
- Classification scheme by Halperin/Hohenberg
 - “Models”, ordered by conserved fields and non-vanishing Poisson brackets
 - φ^4 w. Langevin dynamics: Model A (3D: $z \approx 2.05$, 2D: $z \approx 2.17$)
 - φ^4 w. Hamiltonian dynamics: Model C (3D: $z = 2.17$, 2D: $z = 2$)
 - QCD: Model H $\rightarrow z = 3$ (Son, Stephanov 2004)

Find universal scaling functions, distinguish models A and C!

Figure: 2D Model A spectral functions at $\tau \approx 0$, $p > 0$

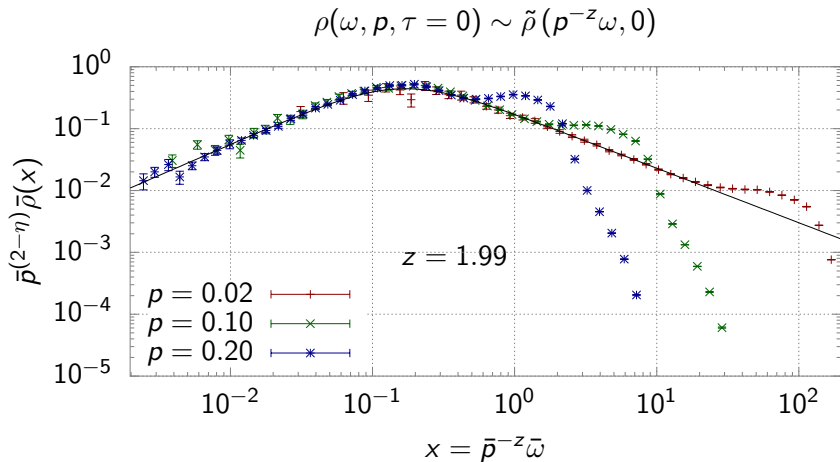
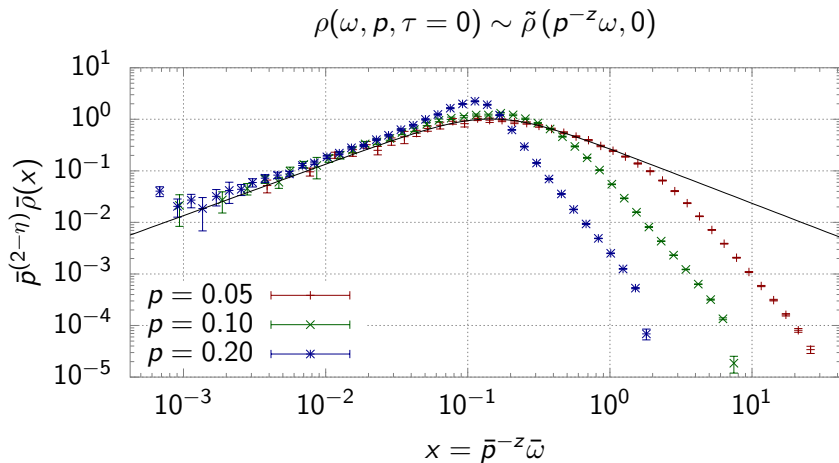
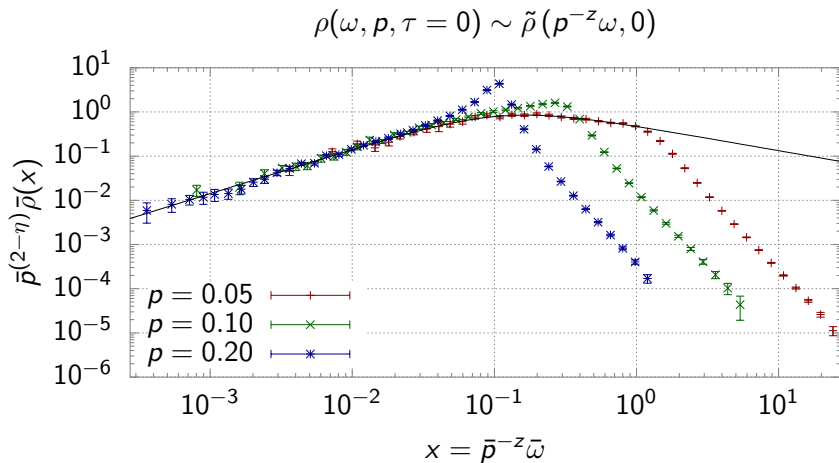


Figure: 2D Model C spectral functions at $\tau \approx 0, p > 0$

Figure: 3D Model A spectral functions at $\tau \approx 0$, $p > 0$

Figure: 3D Model C spectral functions at $\tau \approx 0$, $p > 0$

$$\rho(\omega, p = 0, \tau) \sim \tilde{\rho}(0, \omega/\tau^{\nu z})$$

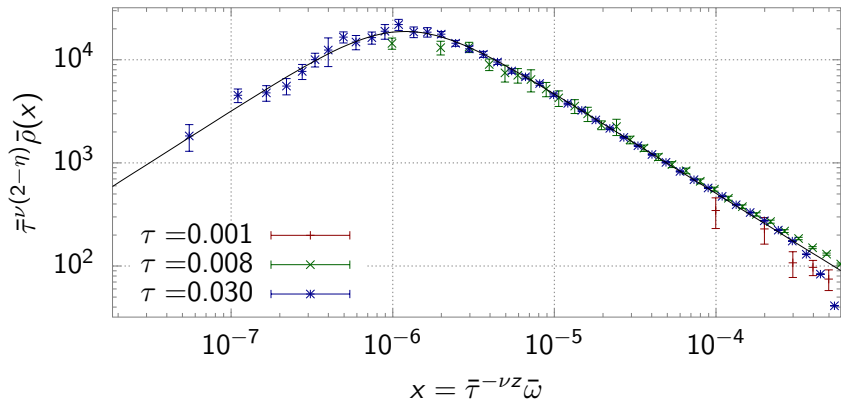


Figure: 2D Model A spectral functions at $\tau > 0, p = 0$

$$\rho(\omega, p = 0, \tau) \sim \tilde{\rho}(0, \omega/\tau^{\nu z})$$

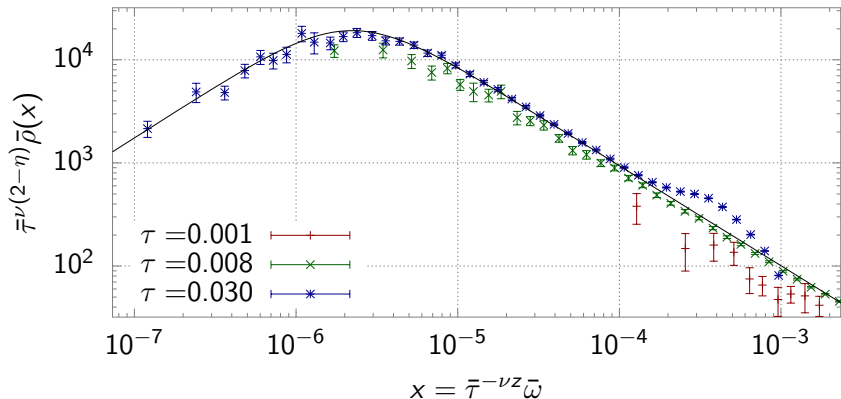


Figure: 2D Model C spectral functions at $\tau > 0, p = 0$

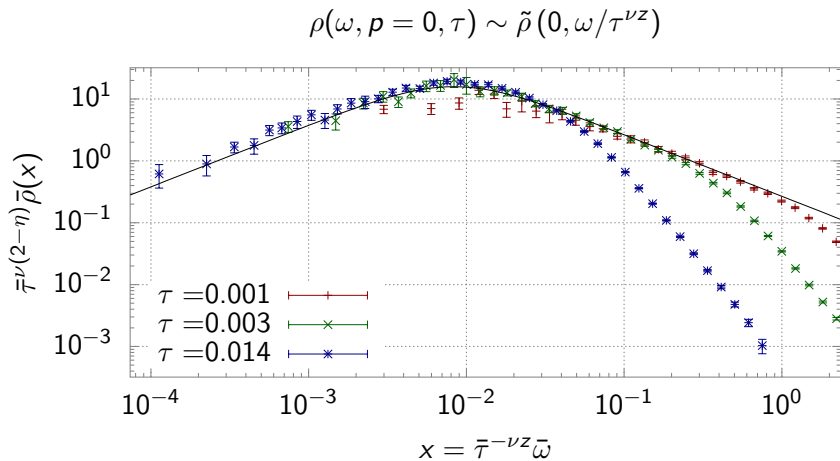
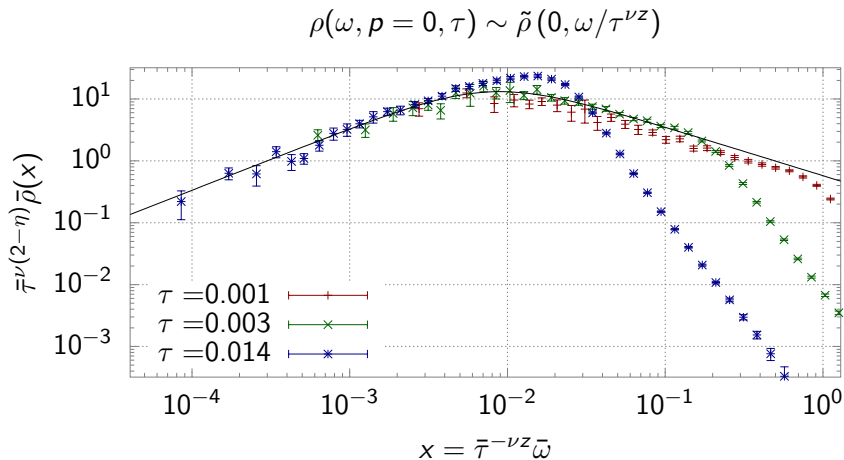


Figure: 3D Model A spectral functions at $\tau > 0, p = 0$

Figure: 3D Model C spectral functions at $\tau > 0, p = 0$

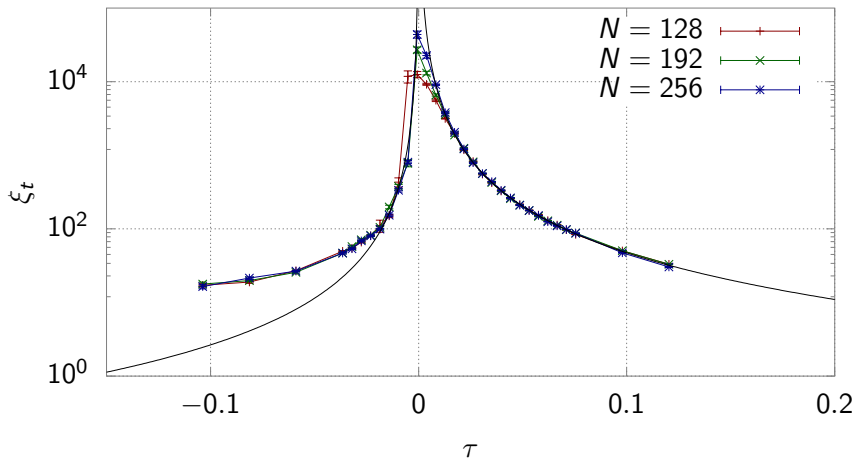


Figure: 2D Model A correlation times with fit function

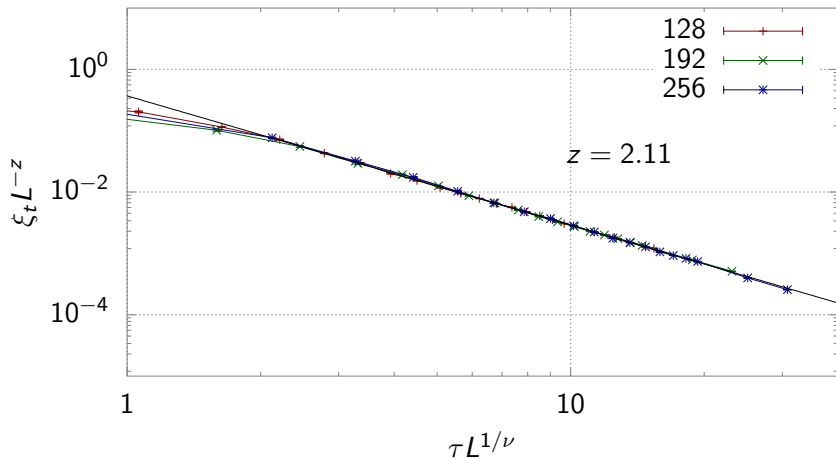


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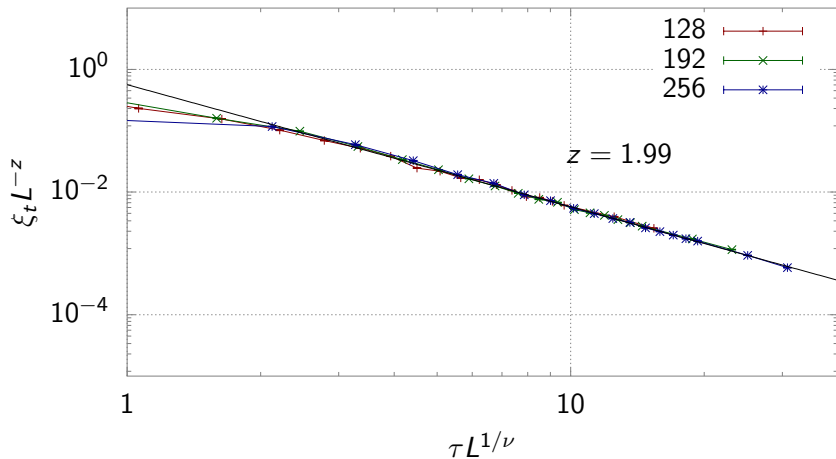


Figure: 2D Model C correlation times with fit function

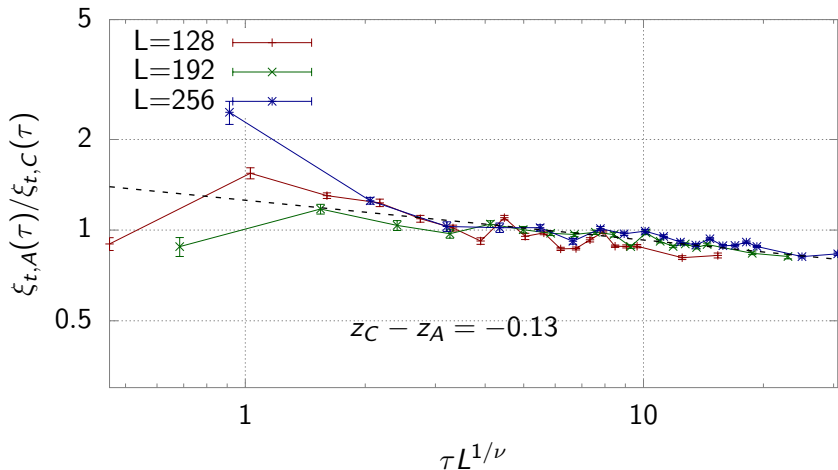


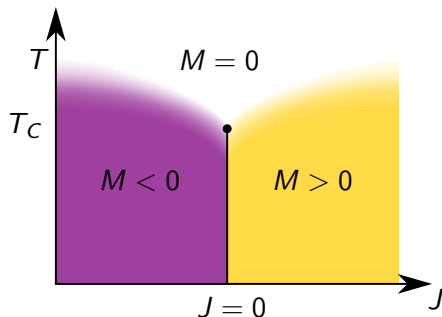
Figure: Ratios of correlation times in 2D. The exponent of the powerlaw corresponds to $\nu(z_C - z_A)$.

- Plausible results for ξ_t , z in power law fits at $\tau = 0$
- Measureable difference $z_C - z_A$
- $d = 3$ has lower amplitudes on ξ_t than $d = 2$
- Model C has stronger non-critical contributions than model A

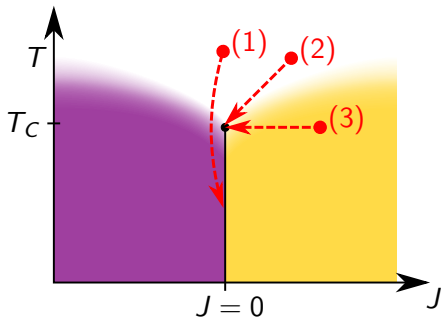
d	Model	$z_{meas.}$	Exp.	MC/scaling
2	A	2.12(1)	2.09(6)	2.1665(12)
2	C	1.98(1)	-	$2 + \frac{\alpha}{\nu} = 2$
3	A	2.02(3)	-	2.05(3)
3	C	2.18(7)	-	$2 + \frac{\alpha}{\nu} = 2.17$

Figure: Comparison of our results to experiment and literature.

- Langevin dynamics couple to heat bath
→ handle on T
 - Vary J over time
- ⇒ Trajectory in phase diagram
- Trajectory near critical point: ξ_t
large, system falls out of equilibrium

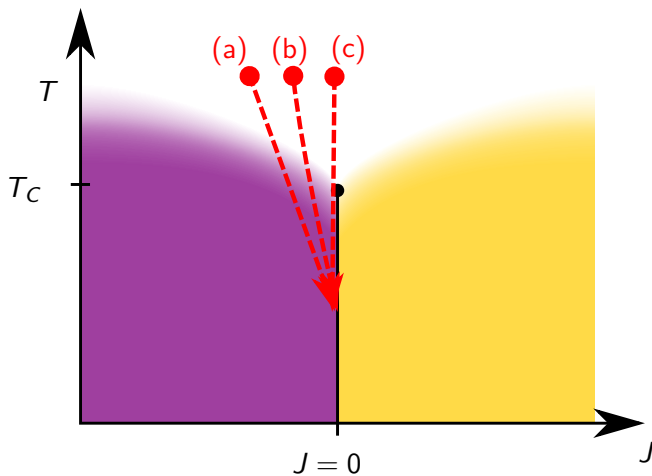


- Thermalize at $T = T_i, J = J_i$,
instant quench to
 $T = T_q, J = J_q$
- Classify non-equilibrium phenomena
 - coarsening, phase ordering (1)
 - initial slip, exponent θ (2)
 - aging (3)



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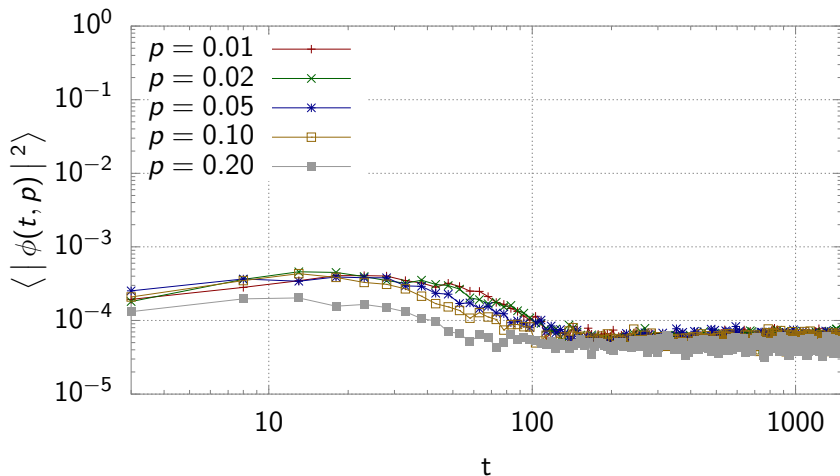


Figure: Order parameter modes fall out of equilibrium

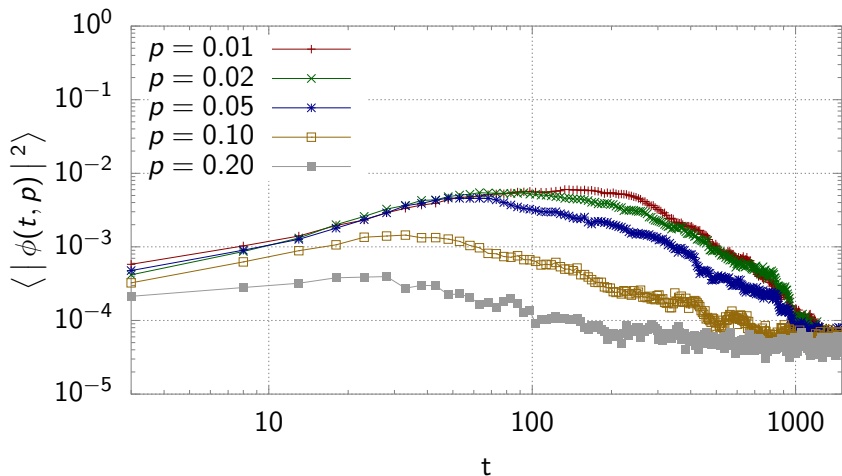


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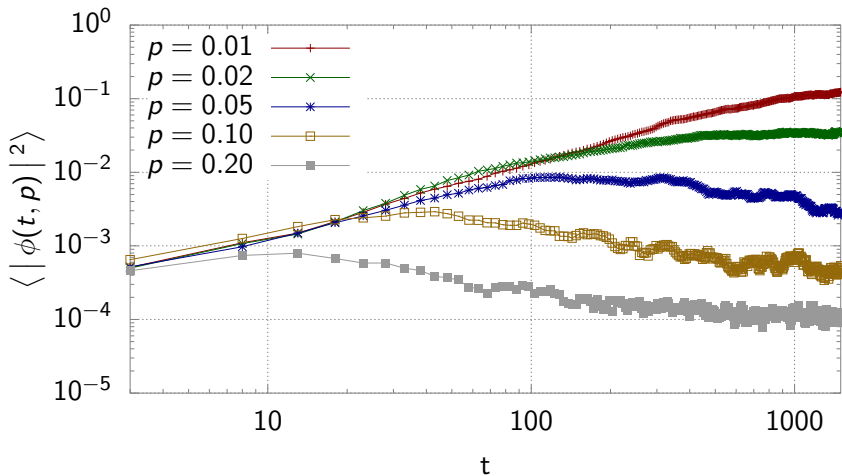
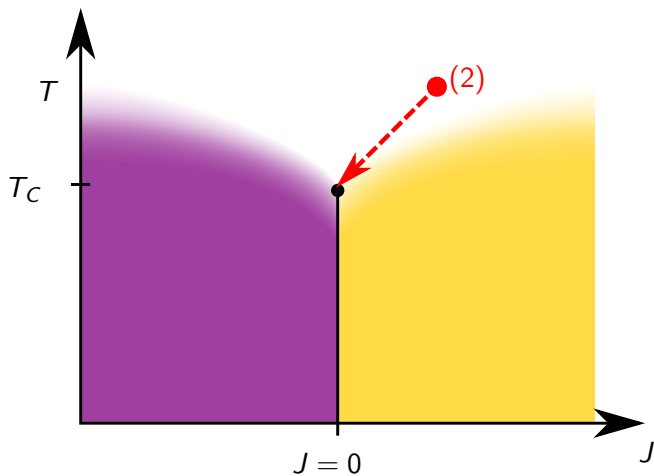


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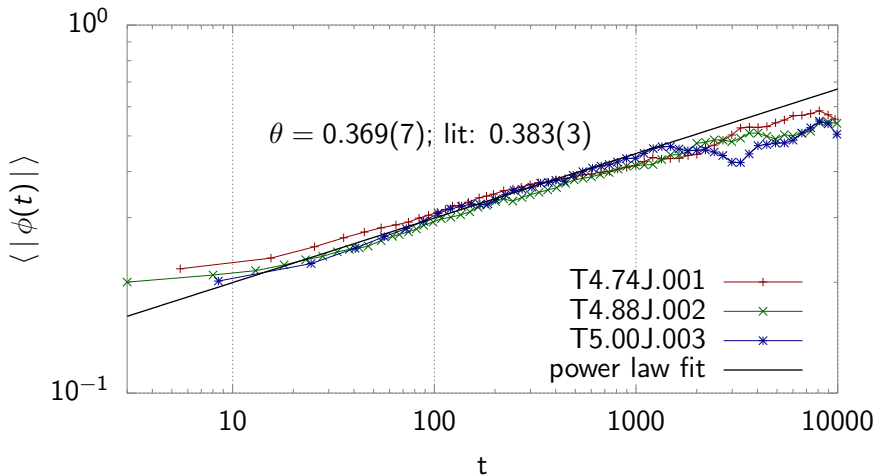
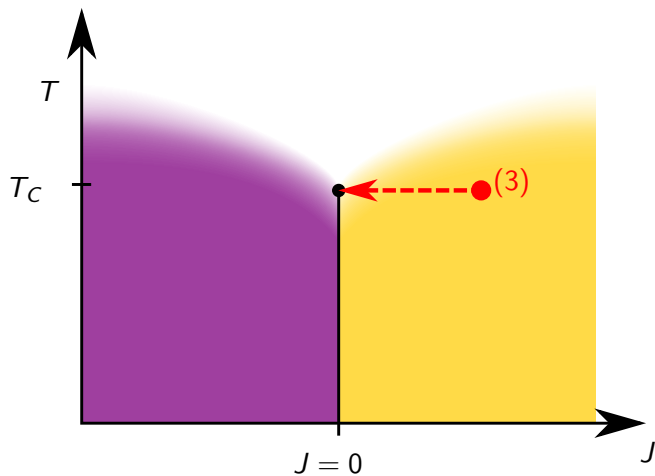


Figure: Algebraic growth, controlled by “initial slip” exponent θ



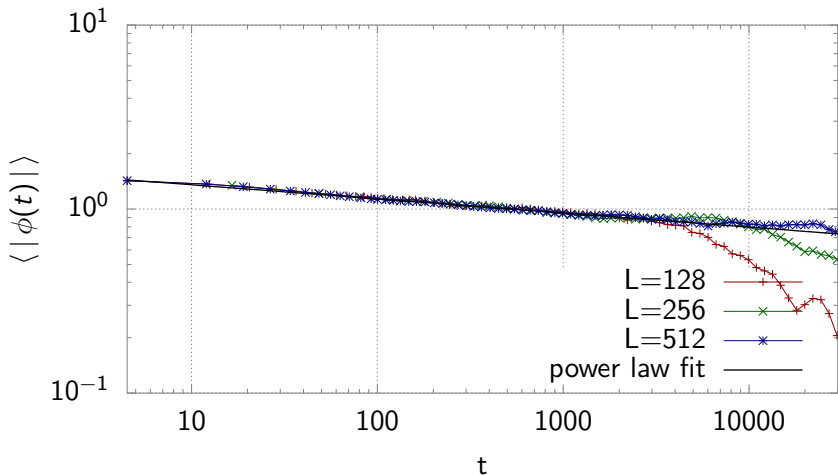
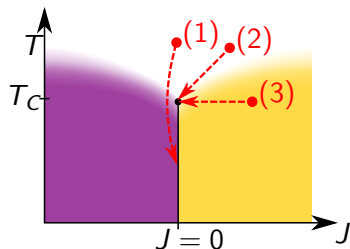


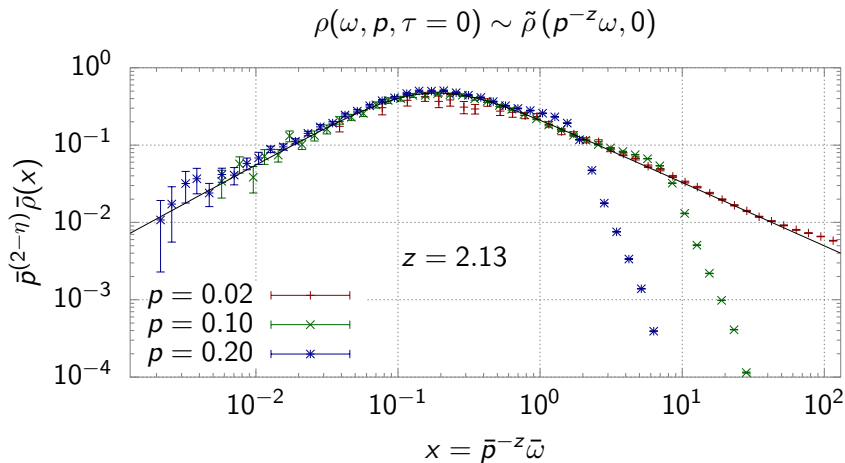
Figure: Algebraic decay, controlled by equilibrium exponents

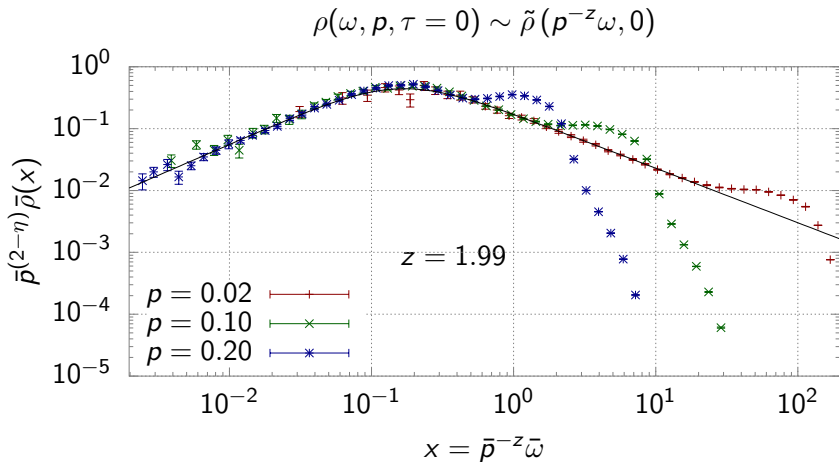
- We can...
 - ... compute spectral functions from lattice data
 - ... extract scaling function of the spectral function
 - ... calculate dynamic critical exponent z , correlation time ξ_t

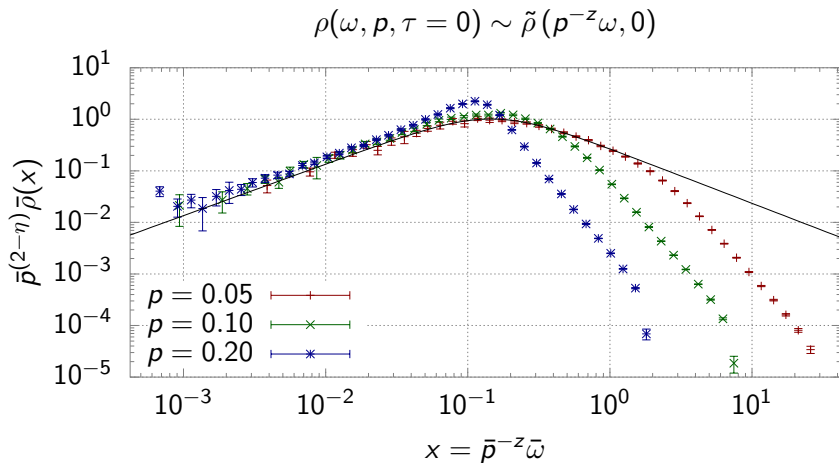
- Where to go from here:
 - Explore non-equilibrium phenomena
 - Apply framework to different models
 - Map results onto QCD



Backup

Figure: 2D Model A spectral functions at $\tau \approx 0, p > 0$

Figure: 2D Model C spectral functions at $\tau \approx 0, p > 0$

Figure: 3D Model A spectral functions at $\tau \approx 0$, $p > 0$

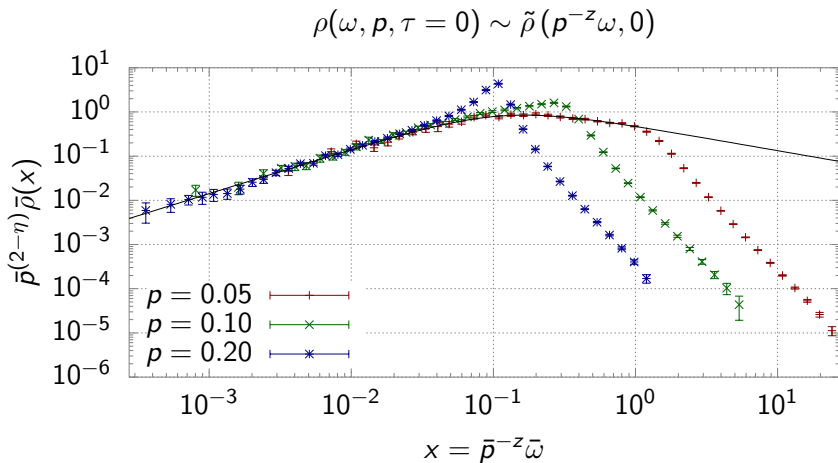


Figure: 3D Model C spectral functions at $\tau \approx 0$, $p > 0$

$$\rho(\omega, p = 0, \tau) \sim \tilde{\rho}(0, \omega/\tau^{\nu z})$$

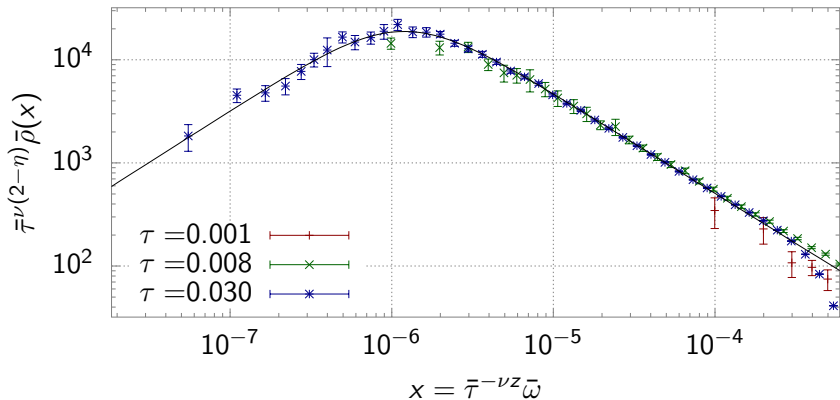


Figure: 2D Model A spectral functions at $\tau > 0, p = 0$

$$\rho(\omega, p = 0, \tau) \sim \tilde{\rho}(0, \omega/\tau^{\nu z})$$

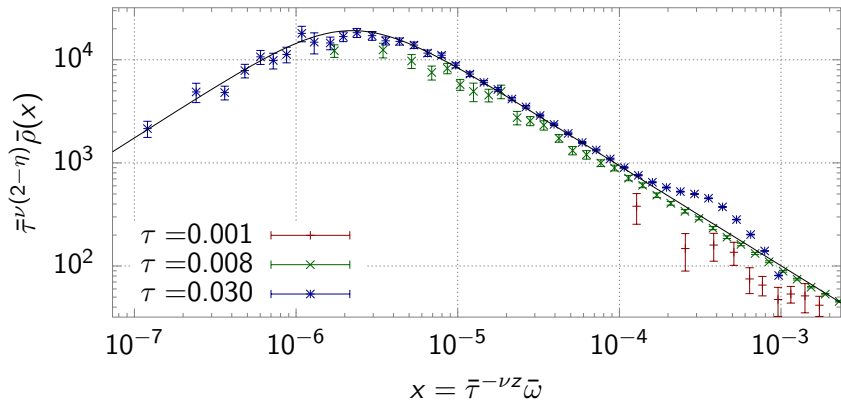
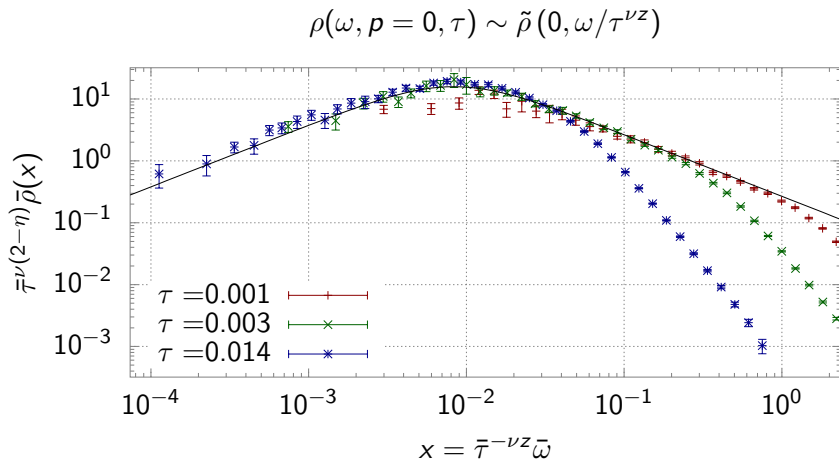
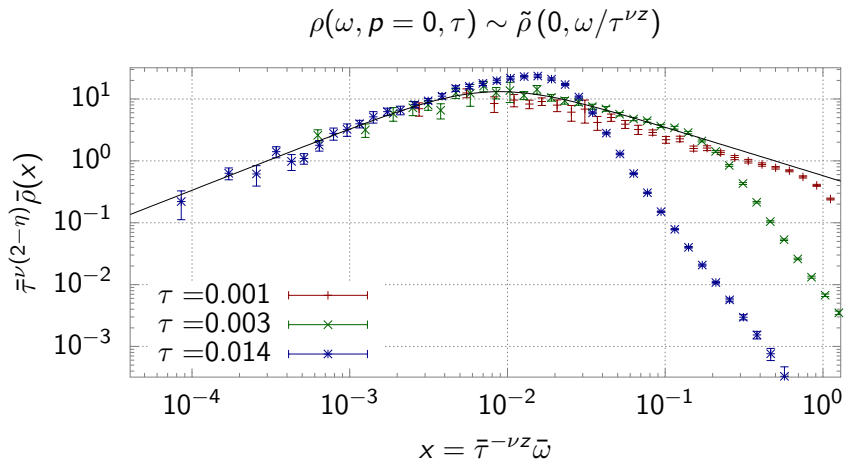


Figure: 2D Model C spectral functions at $\tau > 0, p = 0$

Figure: 3D Model A spectral functions at $\tau > 0, p = 0$

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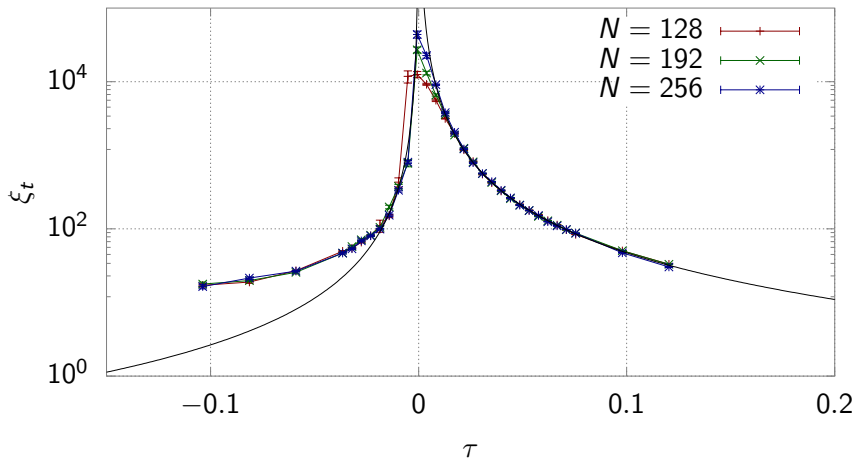


Figure: 2D Model A correlation times with fit function

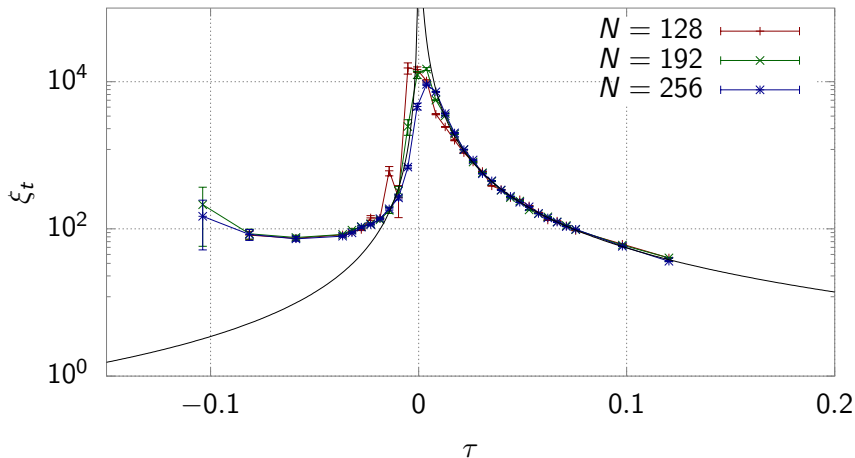


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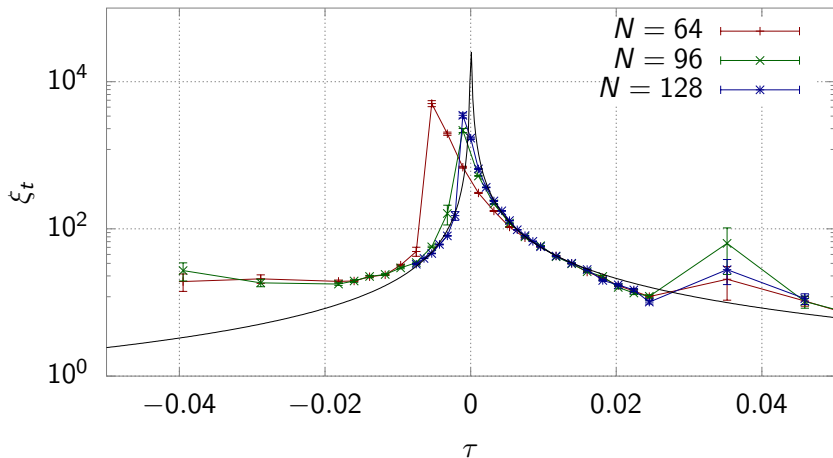


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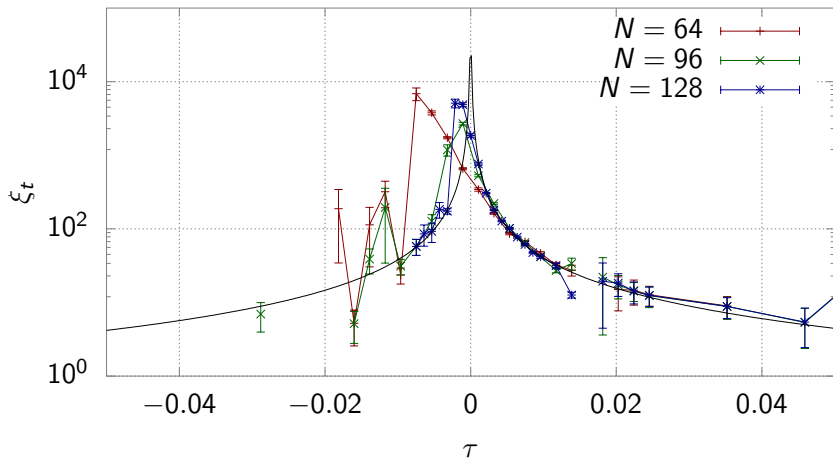


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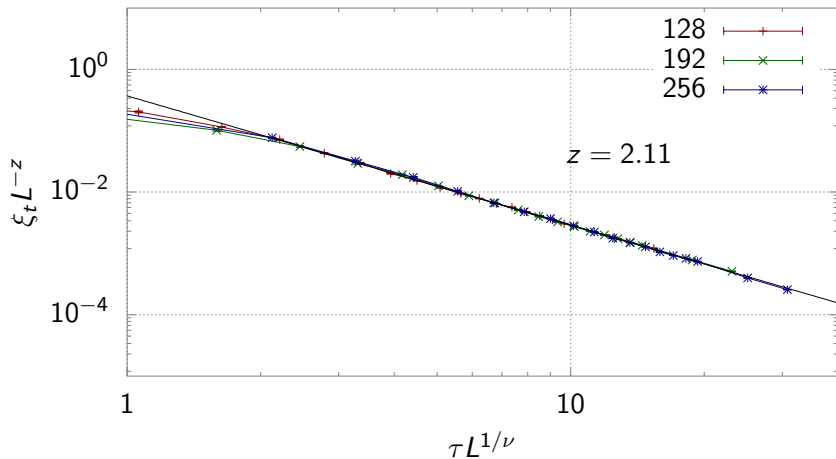


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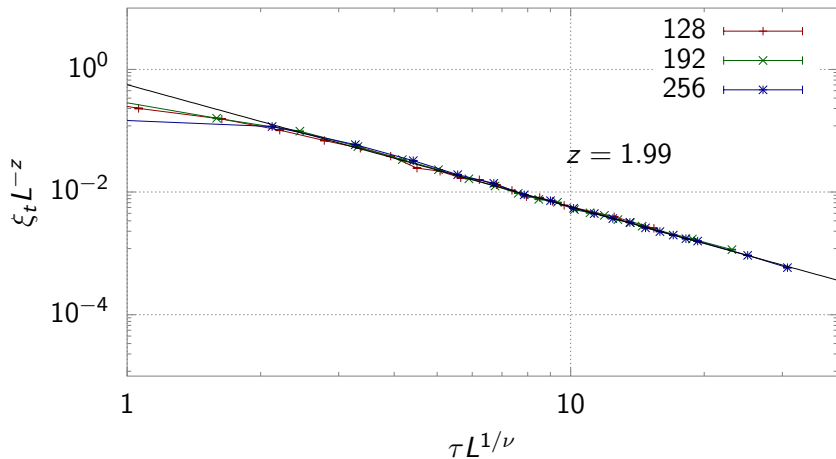


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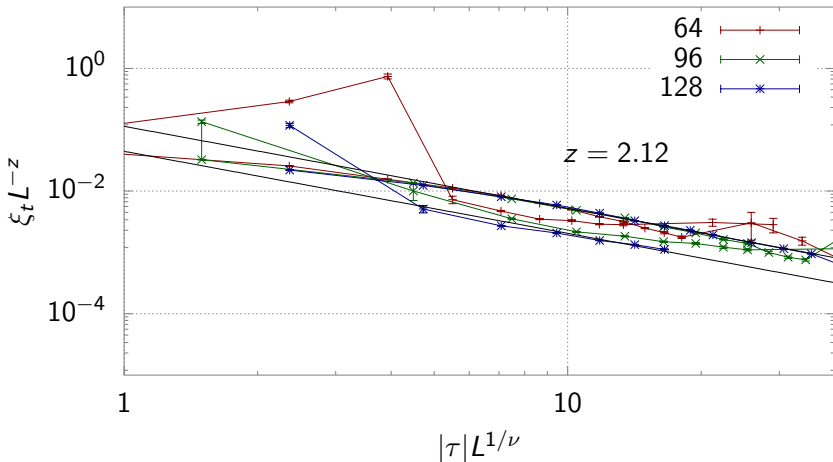


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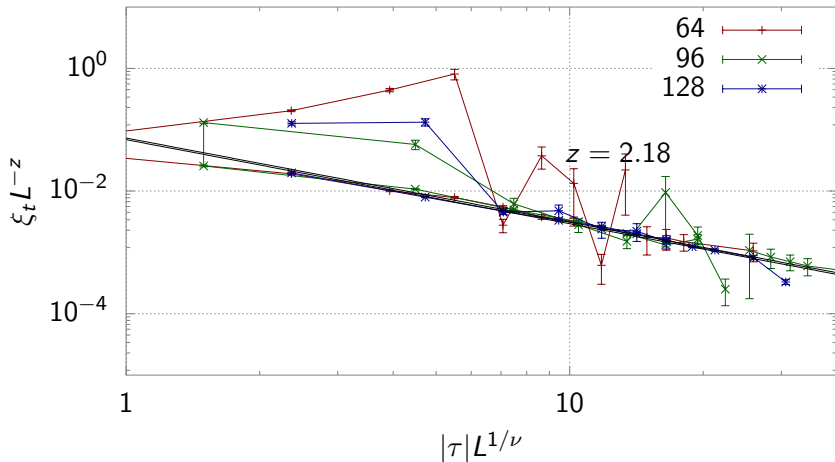


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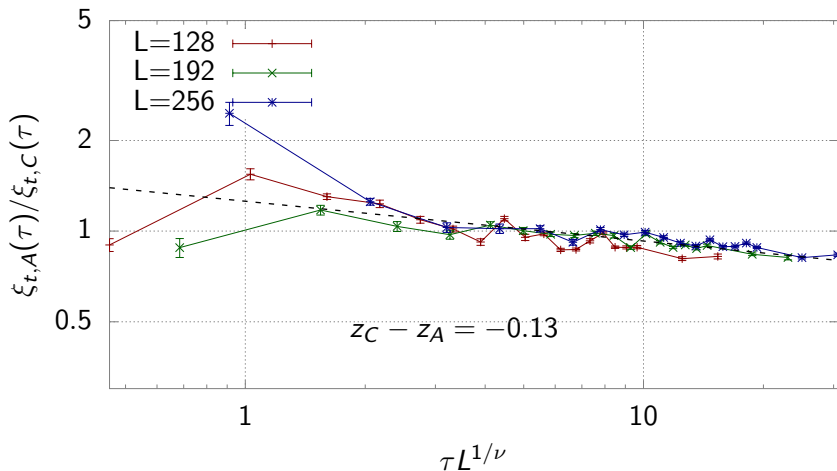


Figure: Ratios of correlation times in 2D. The exponent of the powerlaw corresponds to $\nu(z_C - z_A)$.

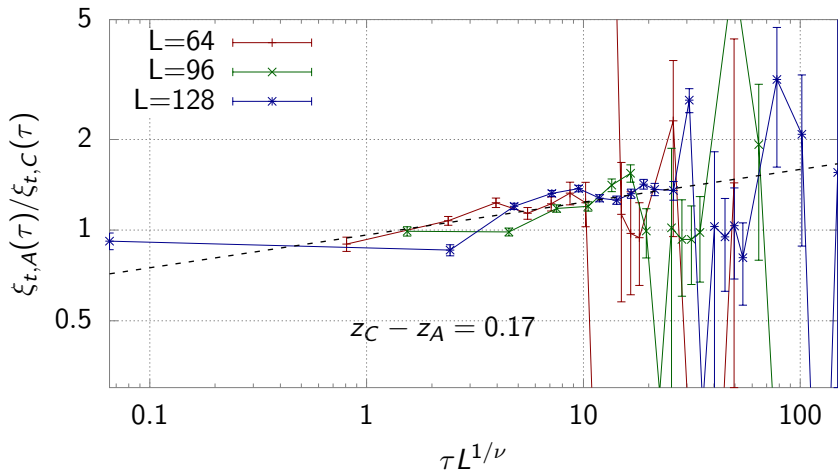


Figure: Ratios of correlation times in 3D. The exponent of the powerlaw corresponds to $\nu(z_C - z_A)$.

- Plausible results for ξ_t , z in power law fits at $\tau = 0$
- Implausibly small results for z in fits at $\tau > 0$
- Measureable difference $z_C - z_A = -0.13(3) < 0$

Model	$z_{meas.}$	Exp.	MC/scaling
A	2.12(1)	2.09(6)	2.1665(12)
C	1.98(1)	-	$2 + \frac{\alpha}{\nu} = 2$

Figure: Comparison of our results to experiment and literature

- Much lower non-universal amplitudes than in 2D
 - power laws harder to fit
 - harder to get good results on z
- Still measurable difference $z_C - z_A = 0.17(9) > 0$

Model	$z_{meas.}$	MC/scaling
A	2.02(3)	2.05(3)
C	2.18(7)	$2 + \frac{\alpha}{\nu} = 2.17$

Figure: Comparison of our results to experiment and literature

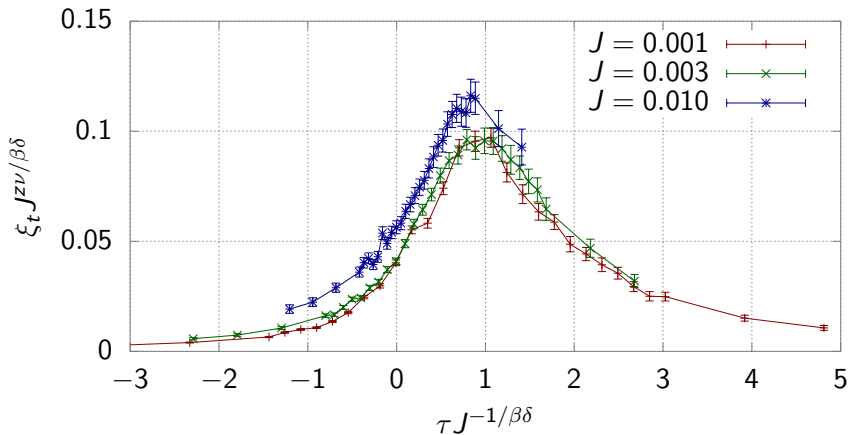


Figure: 2D Model A: Scaling of ξ_t at finite J .