Strangeness in Neutron Stars

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Justus-Liebig-Universität Gießen

December 15, 2021

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A Brief History of the Hyperon Crisis

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Brief History

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• 1932 The neutron is discovered (Chadwick)

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- 1967 Jocelyn Bell & Anthony Hewish discover PSR B1919+21

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In the beginning there was the TOV limit

$$\frac{dp}{dr} = -\frac{G}{r^2} \left(\epsilon(p(r)) + \frac{p}{c^2} \right) \left(M + 4\pi r^3 \frac{p}{c^2} \right) \left(1 - \frac{2GM}{c^2 r} \right)^{-1}$$

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From realistic potentials V_{NN}(r) which are fits to scattering data

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 - Relativistic Models (RMF/RHF)

$$\mathcal{L} = ar{\Psi}_B (i \partial \!\!\!/ - \mathcal{M}_B) \Psi_B + \mathcal{L}_{ ext{mesons}} + \mathcal{L}_{ ext{interactions}}$$

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$$\mathcal{L} = \bar{\Psi}_B (i \partial \!\!\!/^\star - \mathcal{M}_B^\star) \Psi_B + \dots$$

where $M^{\star} = M - g_{\sigma}\bar{\sigma}$ and $p_{0}^{\star} = p_{0} + g_{\omega}\bar{\omega}$



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 Well enough to produce strangeness...

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 Well enough to produce strangeness...

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- PSR J1614-2230
 M = 1.908 ± 0.016 M_☉
- PSR J0348+0432
 M = 2.01 ± 0.04 M_☉
- PSR J0740+6620
 M = 2.08 ± 0.07 M_☉

Hyperon Puzzle in a Nutshell

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Hyperon Puzzle in a Nutshell



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• Modify gravity!

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- Kaon condensation

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Some Solutions?

- Modify gravity!
- Extra hyperon repulsion
- Kaon condensation
- Also the deconfinement question...

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The Quark-Meson Coupling Model

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• The QMC model is one of the relativistic phenomenological models. What we want, fundamentally, is

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- However, we also want to retain some information on the *baryon structure*.
- We choose to model baryon-baryon interactions as a quark-meson interaction in the subhadronic level.



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Bag Model

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where $m^* = m - g^q_\sigma \bar{\sigma}$

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 M_B



 M_B^{\star}



 $M_B^*(\bar{\sigma})$

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• We then calculate the energy of a stationary bag, that is, the mass of the baryon

$$M_B^{\star}(\bar{\sigma}) = M_B - g_{\sigma}^B \bar{\sigma} + \frac{d}{2} (g_{\sigma}^B \bar{\sigma})^2$$

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Some QMC Highlights

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Some QMC Highlights



Figure: Nuclei binding energies from Martinez et al. [2019]

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Some QMC Highlights



Figure: Stellar structure results from Motta et al. [2021] plus recent NICER results such as et al [2021]

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Results

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 $\epsilon(n_n, n_p, n_e)$

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 - Electrically neutral $\rightarrow \sum n_i Q_i = 0$
From Micro to Macro

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 - Energetically stable → Minimize the energy density

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$$\mathcal{L}_{q\delta} = g^{q}_{\delta} \bar{\psi} \boldsymbol{\tau} \cdot \boldsymbol{\delta} \psi$$

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 $M_B^{\star}(\sigma, \delta)$

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$$\mathcal{L}_{q\delta} = g^{q}_{\delta} \bar{\psi} \boldsymbol{\tau} \cdot \boldsymbol{\delta} \psi$$

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$$M_B^{\star}(\sigma,\delta) = M_B - \left(g_{\sigma} + \frac{d}{2}g_{\sigma}^2\sigma\right)\sigma - \left(g_{\delta} - d_2g_{\sigma}g_{\delta}\sigma\right)\mathbf{t}\cdot\boldsymbol{\delta}$$

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• This complicates things. And we **do not expect a big change on the mass** of the neutron star.

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- Does it change other quantities as well? E.g. Tidal deformability?

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$$\Lambda \propto \frac{Q}{{\cal E}}$$

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• What about the Moment of Inertia?

Moment of Inertia



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Tidal Deformability



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• What about the crust of the NS?

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• What about the crust of the NS? Does it affect the Moment of Inertia?

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• What about the crust of the NS? Does it affect the Moment of Inertia?



• What about the crust of the NS? Does it affect the Tidal Deformability?

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- We know from previous studies (Zeno Kordov) that in the standard QMC model they do not appear...
- In particular, what is the δ meson influence?

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• So how do we approach the problem?

 $\epsilon(n_n, n_p, n_e, n_\mu, n_\Lambda, n_{\Xi^0}, n_{\Xi^-}, n_{\Sigma^+}, n_{\Sigma^-}, n_{\Sigma^0}, \dots)$

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 $\epsilon(n_n, n_p, n_e, n_\mu, n_\Lambda, n_{\Xi^\circ}, n_{\Xi^-}, n_{\Sigma^+}, n_{\Sigma^-}, n_{\Sigma^\circ}, \dots)$

• The result of such a minimisation is a series of equilibrium relations for the chemical potential, e.g.

$$\mu_{\rm P} = \mu_{\rm P} + \mu_{\rm e}$$

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 $\epsilon(n_n, n_p, n_e, n_\mu, n_\Lambda, n_{\Xi^0}, n_{\Xi^-}, n_{\Sigma^+}, n_{\Sigma^-}, n_{\Sigma^0}, \dots)$

• The result of such a minimisation is a series of equilibrium relations for the chemical potential, e.g.

$$\mu_{
m P} = \mu_{
m P} + \mu_{
m e}$$

• For the Delta, the relation would be

$$\mu_{\Delta^-} = \mu_n + \mu_e$$

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• So how do we approach the problem?

 $\epsilon(n_n, n_p, n_e, n_\mu, n_\Lambda, n_{\Xi^0}, n_{\Xi^-}, n_{\Sigma^+}, n_{\Sigma^-}, n_{\Sigma^0}, \dots)$

• The result of such a minimisation is a series of equilibrium relations for the chemical potential, e.g.

$$\mu_{
m P} = \mu_{
m P} + \mu_{
m e}$$

• For the Delta, the relation would be

$$\mu_{\Delta^-} = \mu_{\rm D} + \mu_{\rm e}$$

• We can calculate μ_{Δ^-} by

$$\mathcal{M}_{\Delta^{-}} + \sum_{\varphi, B} \left| \bigvee_{\varphi} \right\rangle^{B}$$

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The Nature of Strangeness Puzzle

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• The QMC and other models have shown good agreement with masses and radii *including hyperons*

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- The Hyperon puzzle as it was first conceived of doesn't seem like too much of a puzzle anymore

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- The QMC and other models have shown good agreement with masses and radii *including hyperons*
- The Hyperon puzzle as it was first conceived of doesn't seem like too much of a puzzle anymore
- However, one puzzle to be addressed is that of deconfinement.
- One interesting approach is to try the inverse problem, get NS constraints and from there infer what the EoS is. See Annala et al. [2020]

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Figure: Results from Annala et al. [2020]



Figure: Motta et al. [2021] comparing QMC with Annala et al. [2020]

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Conclusions

- The QMC model is in agreement with data
- There isn't much of a Hyperon puzzle anymore
- We still cannot discriminate between QM and Hyperonic matter



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